

Internal wave generation by tidal flow over a two-dimensional ridge: energy flux asymmetries induced by a steady surface trapped current

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(Received 20 December 2016; revised 1 September 2017; accepted 31 October 2017;
first published online 11 December 2017)

The effects of a surface trapped steady background current on internal waves generated by tidal currents oscillating over a small symmetric ridge are investigated using a two-dimensional primitive equation model. A rigid lid is used with a linearly stratified fluid and the effects of rotation are not considered. We consider uni-directional background currents $\bar{U}(z) \geq 0$ confined to a surface layer lying well above the ridge. The current introduces asymmetries in the generated wave field. For sufficiently narrow ridges the upstream energy flux is larger than the downstream flux while the opposite is the case for sufficiently wide ridges. The total energy flux radiating away from the ridge is not significantly affected by the current. Mean second-order currents and pressure fields are shown to make important contributions to the total energy flux. A first-order linear theory, valid for a general stratification and surface current, which accurately predicts the wave field is also developed.

Key words: geophysical and geological flows, internal waves

1. Introduction

In this paper we investigate the effects of a steady, surface trapped background current $\bar{U}(z)$ on the two-dimensional generation of oceanic internal waves by tidal currents oscillating over a symmetric ridge in the weakly nonlinear regime. Figure 1(a) shows the wave-induced horizontal currents after 15 tidal periods from a fully nonlinear numerical simulation (details of the simulations are provided in §2). The background current lies above $z/H = -0.3$, z and H being the vertical coordinate and deep water depth respectively, and is directed to the right. Obvious asymmetries in the wave field are apparent. Mode-one waves have propagated outside the region shown. Mode-two waves have reached $x/H \approx -80$ and 120 in the upstream and downstream directions. Internal wave beams are much stronger in the upstream direction and there is a fan-like structure near the surface downstream of the ridge (x/H between 0 and 80). Figure 1(b) shows the wave-induced horizontal baroclinic currents predicted by linear theory (§4) in which the contributions for each mode are restricted to lie within the distance they would propagate in 15 tidal periods at

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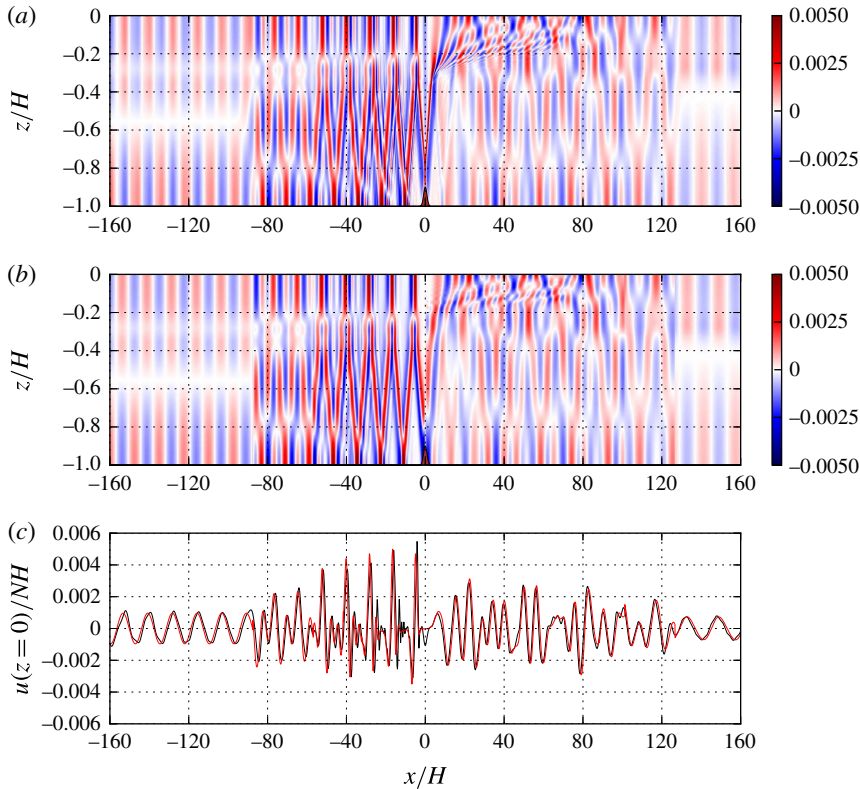


FIGURE 1. Comparison of wave-induced dimensionless horizontal current fields $U/(NH)$ at $t = 15T$ from a numerical simulation (a) and theory (b). Here N is the constant buoyancy frequency and H the deep water depth. Panel (c) compares the wave-induced horizontal currents at the surface (black, simulation; red, theory). Narrowest ridge from set 4 (see table 1): $h_0/H = 0.1$, $a/H = 1.2$, $(U_s/(NH), z_s/H, d_s/H) = (0.1, -0.3, 0.08)$.

their group velocity. The theory predicts the wave field quite well although there are differences. Figure 1(c) compares the simulated and theoretical wave-induced surface currents showing excellent agreement.

Upstream of the ridge the wavelength of the beams is approximately 11.4 versus the mode-one wavelength of 12.5. The strong horizontal currents where the beam reaches the surface are clearly detectable in figure 1(c). Because the wavelengths of the different modes are non-commensurable due to the background current, the internal wave beams that form lose their coherence with distance from the ridge. This also explains the difference in beam wavelength from the mode-one wavelength.

Figure 2 shows the mean wave-induced horizontal currents for a slightly wider ridge. The structure of the mean currents upstream and downstream of the ridge are very different. The quasi-periodic pattern upstream (left) of the ridge is reflected in the energy flux contributions (see § 6). The mean current upstream of the ridge is large at the surface, where it is always positive, at the base of the shear layer, and where beams reflect off the bottom. At the base of the shear layer there are alternating patches of positive and negative mean currents with the latter higher up than the former. This is consistent with beams reaching the base of the current and exerting a stress which raises the base of the current slightly, hence reducing the

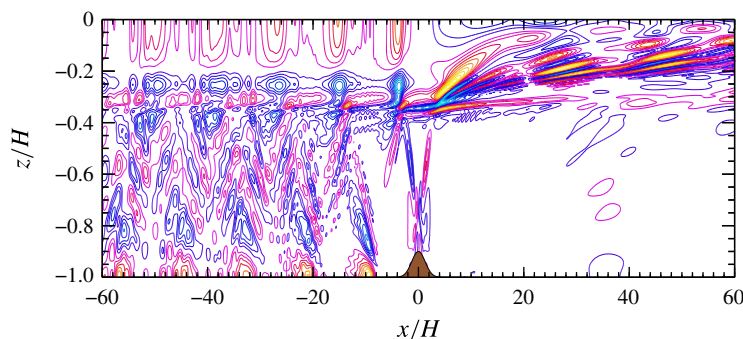


FIGURE 2. Mean horizontal velocity perturbation (dimensionless). Similar case to that depicted in figure 1 (slightly wider ridge from set 4 with $a/H = 1.8$: see table 1). Averaging done over $t \in [36T, 40T]$, where T is the tidal period, in a Lagrangian reference frame moving with the barotropic tidal currents. Green–blue colours are negative currents, red–orange are positive currents. Contour interval 10^{-5} .

mean current, and lowering the base of the current slightly between the beams, hence increasing the mean current. To the left of the ridge the maximum and minimum mean currents are approximately 10^{-4} , approximately a factor of 50 less than the currents associated with the waves of tidal frequency (see figure 1). In dimensional terms they are approximately 0.4 mm s^{-1} . Downstream of the ridge the mean currents are about twice as strong.

Figures 1 and 2 show significant asymmetries in the wave field upstream and downstream of the ridge. Associated with this are asymmetries in the energy fluxes which is the primary focus of this paper. Somewhat surprisingly, in the parameter regime we consider, for wide ridges the upstream energy flux is smaller than the downstream energy flux while the opposite is the case for narrow ridges. In the presence of a strong background current the kinetic energy flux makes an important contribution to the total flux and fluxes associated with the mean wave-induced fields are also significant.

Internal wave generation by tide–topography interactions has received a lot of attention over the past couple of decades. This process is estimated to transfer energy from the barotropic tide to internal tides at a rate of approximately 1 TW and accounts for approximately 25 %–30 % of the barotropic tidal energy dissipation (Egbert & Ray 2001; Nikurashin & Ferrari 2013; Waterhouse *et al.* 2014). It is also the source of approximately 50 % of the internal wave energy in the ocean, approximately of equal importance to wind forcing. Accurate modelling or parametrizations of this process are necessary for modelling global barotropic tides (Lyard *et al.* 2006; Buijsman *et al.* 2016) and for taking proper account of mixing by internal waves which is believed to be essential for the large scale overturning circulation in the ocean (Munk & Wunsch 1998; Wunsch & Ferrari 2004; Nikurashin & Ferrari 2013). Energy transferred to the internal wave field is partly dissipated near the bottom above rough topography (Polzin *et al.* 1997; Waterhouse *et al.* 2014; Lefauve, Muller & Melet 2015; Ferrari *et al.* 2016) but can propagate large distances before dissipating when the waves are primarily generated by isolated bathymetric features (Falahat *et al.* 2014). Internal tide generation in the deep ocean was reviewed by Garrett & Kunze (2007).

Geostrophic currents also interact with bathymetry to generate internal lee waves injecting approximately 0.2–0.4 TW into the internal wave field (Scott *et al.* 2011;

Nikurashin & Ferrari (2013) predominantly in the Southern Ocean where the interaction of the Antarctic Circumpolar Current with bathymetry is an important generator of internal waves.

Considering the ubiquitous presence of large scale currents and eddies in the ocean many tide–topography interactions will inevitably involve background currents although there have been few studies of this process. Pickering *et al.* (2015) considered the effects of mean flows on wave generation and energy fluxes in Luzon Strait during periods when the Kuroshio Current passes through the strait. The focus of their study was on energy fluxes at the generation site, not on energy fluxes in the far field. The Indonesian Throughflow passing through Lombok Strait (Aiki, Matthews & Lamb 2011; Matthews *et al.* 2011) and the South Equatorial Current passing over the Mascarene Plateau (da Silva, New & Magalhaes 2011) have been shown to affect the generation of internal solitary waves. Another example includes exchange flows through the Straits of Gibraltar. All of these examples involve large sills which interact directly with the currents.

The effects of a mean flow over ridge-like topography on internal wave generation by tide–topography interactions have not, to our knowledge, been previously considered in numerical or theoretical studies. Here we address this lack by considering the effects of a background current $\bar{U}(z)$ confined to a surface layer lying well above the ridge so that there is no direct interaction between the current and the bathymetry which would result in the co-generation of lee waves. We conduct numerical simulations using a fully nonlinear two-dimensional model. For simplicity our background currents satisfy $\bar{U}(z) \geq 0$ and a linear stratification is used.

A linear theory is also developed to predict the internal waves generated by the tide–topography interactions. Our theory is an extension of that presented by Bell (1975*a,b*) and Khatiwala (2003). Bell (1975*b*) considered an infinitely deep ocean predicting the vertically propagating internal waves generated by weak oscillating currents over a small subcritical isolated topographic feature. Khatiwala (2003) extended this method to a finite depth ocean with a linear stratification. We extend his method to an arbitrary stably stratified shear flow with the restriction that the current lies above the bathymetry. One cost of this extension is that fewer results can be computed analytically. The theory is valid for small amplitude, subcritical bathymetry and small tidal excursion distances. Neither the theory nor the numerical simulations include the effects of rotation as this would require a transverse pressure gradient in geostrophic balance with the background current. While rotation could be included in the numerical simulations, it implies a density field that varies in the cross-current direction, making the problem three-dimensional.

Section 2 describes the numerical model and the simulations. Energy flux, as calculated in the nonlinear, non-hydrostatic numerical simulations, is discussed in § 3 and the linear theory is developed in § 4. Results of convergence and validation tests are given in § 5 while the main results are presented in § 6 followed by the conclusions and summary in § 7.

2. Numerical model and model set-up

We use the two-dimensional non-hydrostatic Internal Gravity Wave model (Lamb 1994, 2007) to solve the incompressible Euler equations under the Boussinesq approximation in a non-rotating reference frame. The model equations are

$$u_t + uu_x + wu_z = -p_x, \quad (2.1a)$$

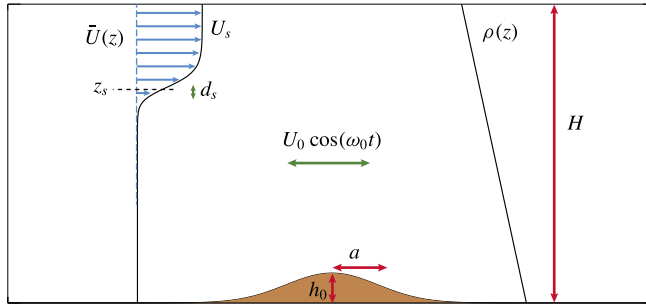


FIGURE 3. Schematic of the numerical set-up. \bar{U} is a steady background current with surface current U_s , base at $z=z_s$ and d_s is a measure of the width of the shear layer. h_0 and a are the ridge amplitude and width parameters, $\bar{\rho}$ the linear background stratification and $U_0 \cos(\omega_0 t)$ the background tidal current. The shape of the current and hill amplitude correspond to set 2 simulations.

$$w_t + uw_x + ww_z = -p_z - \rho g, \quad (2.1b)$$

$$\rho_t + u\rho_x + w\rho_z = 0, \quad (2.1c)$$

$$\nabla \cdot \mathbf{u} = 0. \quad (2.1d)$$

Here $\mathbf{u} = (u, w)$ is the velocity in the vertical xz -plane, ρ is the density and p is the pressure, both of which have been scaled by the reference density ρ_0 . The rigid lid approximation is used with the surface at $z=0$. We consider waves generated by tidal flow over an isolated topographic feature at $z = -H + h(x)$ such that $h(x) \rightarrow 0$ as $x \rightarrow \pm\infty$. The model uses Godunov flux limiting which acts as an implicit large eddy simulation model (Bell, Colella & Glaz 1989). Terrain following coordinates are used which leads to higher vertical resolution in shallower water. In these simulations the time step is fixed.

Figure 3 shows a schematic of the numerical set-up. The simulations were conducted using dimensional variables with the deep ocean in mind. We used a fluid of depth $H = 5000$ m in the far field and Gaussian ridges

$$h(x) = h_0 e^{-(x/a)^2}, \quad (2.2)$$

which have a maximum slope of

$$s = \sqrt{\frac{2}{e}} \frac{h_0}{a} \approx 0.86 \frac{h_0}{a}. \quad (2.3)$$

To simplify the calculation of available potential energy for the theoretical analysis, we restrict ourselves to a uniform stratification with buoyancy frequency $N = 10^{-3} \text{ s}^{-1}$ for which an analytical expression for the available potential energy exists.

We consider background currents of the form

$$\bar{U}(z) = \frac{U_s}{4} \left(1 + \tanh \left(\frac{z - z_s}{d_s} \right) \right)^2, \quad (2.4)$$

with z_s and d_s chosen so that the current lies above the ridge and U_s is the current at the surface. The minimum Richardson number is

$$Ri = \frac{N^2}{\max(\bar{U}')^2} = \frac{d_s^2}{U_s^2} \left(\frac{27}{16} \right)^2 \times 10^{-6}, \quad (2.5)$$

which is above $1/4$ in all simulations so that the background state is stable. The model is forced by specifying u_t at the left boundary chosen to drive a barotropic tidal current

$$\bar{U}_b(t) = U_0 \cos(\omega_0 t) \quad (2.6)$$

far from the ridge and, consistent with the forcing, the simulations are initialized at peak rightward flow with initial currents

$$\left. \begin{aligned} u &= U_0 \frac{H}{H - h(x)} + \bar{U}(z), \\ w &= -U_0 \frac{H}{(H - h(x))^2} h'(x)z. \end{aligned} \right\} \quad (2.7)$$

The initial flat isopycnals are near their mean position. The initial flow is divergence free but has some weak vorticity. Test cases showed similar results if the simulations were started from rest with $u = U_0 \sin(\omega_0 t)$ at the left boundary. We use an M_2 tidal frequency with a period of $T = 12.42$ h (44 712 s) for which the frequency is $\omega_0 \approx 1.4053 \text{ s}^{-1}$.

The criticality of the slope, defined as the ratio of the bottom slope to the slope of an internal wave beam of frequency ω_0 , has a maximum value of

$$\gamma = s \frac{\sqrt{N^2 - \omega_0^2}}{\omega_0} \approx 6.06 \frac{h_0}{a}. \quad (2.8)$$

Because the length of the computational domain is chosen to be long enough so that no waves reach either lateral boundary the problem is defined by nine dimensional parameters: three associated with the geometry of the domain, (H , h_0 , a); three associated with the background current, (U_s , z_s , d_s); two associated with the tidal current, (U_0 , ω_0); and N , the buoyancy frequency of the linear background stratification. There are seven dimensionless parameters. Using the ocean depth $H = 5000$ m and the inverse buoyancy frequency $N^{-1} = 1000$ s as the length and time scales, the dimensionless parameters can be chosen as a/H , h_0/H , d_s/H , z_s/H , U_s/NH , ω_0/N and U_0/NH . It is impossible to explore all of parameter space. For comparisons with the linear theory we consider the near-linear limit by considering subcritical slopes and by keeping the dimensionless ridge amplitude h_0/H and tidal excursion distance $(U_0/\omega_0)/H$ small. We focus on the effects of varying the properties of the background current and the width of the ridge which affects the modal composition of the wave field. We also take $\omega_0/N = 0.14$ so we are in the near-hydrostatic limit. The water depth H , the tidal current parameters U_0 and ω_0 and the buoyancy frequency N are fixed.

Eight simulation sets were done between which the ridge amplitude and the parameters of the background current were varied. For each set 11 ridge widths were used, namely $a/H = 1.2, 1.4, 1.6, 1.8, 2.4, 3.0, 3.6, 4.2, 4.8, 5.4$ and 6.0 . Parameter values for each set are provided in table 1 along with some key non-dimensional parameters, including the Froude number $F_r = U_0/c_1$, which is the ratio of the maximum current to the linear mode-one long wave phase speed in the absence of a background current ($c_1 = \sqrt{N^2 - \omega_0^2}H/\pi \approx NH/\pi$), and the maximum slope criticality, which occurs for the narrowest ridge. The minimum value of γ in each set occurs for the widest ridge and is one fifth of the maximum value. The dimensionless tidal excursion distance is 0.071 in all cases.

Results are presented in non-dimensional terms using the water depth H as the length scale, the inverse buoyancy frequency N^{-1} as the time scale and NH as the velocity scale. Times are reported in tidal periods.

Set	h_0 (m)	U_s (m s ⁻¹)	z_s (m)	d_s (m)	$\frac{h_0}{H}$	$\max\{\gamma\}$	$\frac{z_s}{H}$	$\frac{d_s}{H}$	Ri	$\frac{U_s}{NH}$	$F_r = \frac{U_s}{c_1}$
1	250	0.5	−1500	300	0.05	0.25	−0.3	0.06	1.03	0.1	0.32
2	500	—	−1500	300	0.1	0.50	—	—	—	—	—
3	1000	—	−1500	300	0.2	1.01	—	—	—	—	—
4	500	—	−1500	400	0.1	0.50	—	0.08	1.82	—	—
5	500	—	−2500	300	—	—	−0.5	0.06	—	—	—
6	500	0.2	−1500	400	—	—	−0.3	0.08	6.41	0.04	0.13
7	500	0.8	−1500	300	—	—	—	0.06	0.40	0.16	0.51
8	500	—	−2500	300	—	—	−0.5	—	—	—	—

TABLE 1. Parameters for the simulation sets. For each set of runs the ridge width varies over $a/H \in [1.2, 1.4, 1.6, 1.8, 2.4, 3.0, 3.6, 4.2, 4.8, 5.4, 6]$. The maximum slope criticality for each set, $\max\{\gamma\}$, is for $a/H = 1.2$. $c_1 = H\sqrt{N^2 - \omega_0^2}/\pi \approx NH/\pi$ is the mode-one phase speed in the absence of the background current. The dashes indicate that the value is the same as that directly above.

3. Energy conservation and energy flux

In the following we work in an accelerating reference frame moving with the far-field barotropic tidal current $\bar{U}_b(t)$. In this reference frame the horizontal coordinate is

$$\xi = x - \int_0^t \bar{U}_b(t') dt' = x - \frac{U_0}{\omega_0} \sin(\omega_0 t) \quad (3.1)$$

and the horizontal current is

$$\tilde{u}(\xi, z, t) = u(x(\xi, t), z, t) - \bar{U}_b(t). \quad (3.2)$$

We define a new pressure term via

$$\tilde{p} = p(x(\xi, t), z, t) + \frac{d\bar{U}_b}{dt} \xi. \quad (3.3)$$

This removes the part of the horizontal pressure gradient which drives the barotropic tide in the far field. More generally, let tildes denote functions in the new reference frame, i.e. $\tilde{w}(\xi, z, t) = w(x(\xi, t), z, t)$, etc. Under this change of variables the governing equations in the new reference frame are identical to (2.1a)–(2.1d) except for the addition of tildes to all dependent variables and with derivatives with respect to x replaced by derivatives with respect to ξ . The lower boundary is now moving and is at

$$z = -H + \tilde{h}(\xi, t) = -H + h(x(\xi, t)) = -H + h\left(\xi + \int_0^t \bar{U}_b(t') dt'\right) \quad (3.4)$$

and the bottom boundary condition is

$$\tilde{\mathbf{u}} \cdot \hat{\mathbf{n}} = -(\bar{U}_b(t), 0) \cdot \hat{\mathbf{n}} \quad \text{at } z = -H + \tilde{h}(\xi, t). \quad (3.5)$$

Henceforth we will drop the tildes except for \tilde{h} as both $\tilde{h}(\xi, t)$ and $h(x)$ will be used.

Neglecting viscous and diffusive effects the energy equation is

$$\frac{\partial}{\partial t}(E_k + E_a) + \nabla \cdot (\mathbf{u}(E_k + E_a + p_d)) = 0, \quad (3.6)$$

where p_d is the pressure perturbation (or disturbance) relative to the hydrostatic pressure of the undisturbed flow $\bar{p}(z)$,

$$E_k = \frac{1}{2}(u^2 + w^2) \quad (3.7)$$

is the kinetic energy density (all energies are per unit mass) and

$$E_a(\xi, z, t) = g \int_z^{z^*(\xi, z, t)} (\bar{\rho}(s) - \rho(\xi, z, t)) ds \quad (3.8)$$

is the available potential energy density (Holliday & McIntyre 1981; Shepherd 1993). Here $\bar{\rho}(z)$ is the reference density and $z^*(\xi, z, t)$ is the height of the fluid particle at (ξ, z, t) in the reference stratification (Scotti, Beardsley & Butman 2006; Lamb 2007, 2008; Lamb & Nguyen 2009). We use the background stratification as the reference density which is appropriate for calculating the available potential energy in an infinitely long domain (Lamb 2008). It avoids sorting the density field which is computationally expensive. Since the background stratification is uniform, the available potential energy density has a simple analytic expression, namely

$$E_a = \frac{1}{2} \frac{g^2 \rho'^2}{N^2}, \quad (3.9)$$

where ρ' is the density perturbation.

Let

$$u = \bar{U}(z) + u', \quad (3.10)$$

where u' is the horizontal velocity perturbation. Far from the ridge u' is the wave-induced current; however, over the ridge it includes a barotropic contribution due to the constriction of the tidal flow over the ridge.

We split the kinetic energy density into three terms

$$E_k = E_{k0} + E_{k1} + E_{k2}, \quad (3.11)$$

where

$$\left. \begin{aligned} E_{k0} &= \frac{1}{2} \bar{U}^2, \\ E_{k1} &= \bar{U} u', \\ E_{k2} &= \frac{1}{2} (u'^2 + w^2) \end{aligned} \right\} \quad (3.12)$$

are the contributions to the kinetic energy density which are of order zero, one and two in the perturbation velocities. We will refer to

$$E_{kp} = E_{k1} + E_{k2}, \quad (3.13)$$

as the perturbation kinetic energy density. Since \bar{U} is independent of time the energy equation can be written as

$$\frac{\partial}{\partial t}(E_{kp} + E_a) + \nabla \cdot (\mathbf{u}(E_k + E_a + p_d)) = 0. \quad (3.14)$$

The perturbation kinetic energy can be negative if E_{k1} dominates and is negative (Lamb 2010). We do not work with the positive definite wave energy $E_{k2} + E_a$ as this is not a conserved quantity. Use of wave energy, which is the energy associated with the wave in a reference frame moving with the background current, has proved to be a useful quantity in the context of wave packets propagating through a slowly varying background field. As a wave packet moves through a sheared background current wave energy is exchanged with the slowly varying background current and wave action is conserved (Broutman, Rottman & Eckermann 2004). In our situation the majority of the energy is in low vertical modes which have wavelengths comparable to the water depth and the background sheared current does not vary slowly compared to this length scale.

In the following we use bars to indicate energy values integrated over a domain

$$\mathcal{D} = [\xi_l, \xi_r] \times [-H + \tilde{h}(\xi, t), 0], \quad (3.15)$$

which is bounded by stationary lateral boundaries at $\xi = \xi_l$ and ξ_r , by the moving bottom at $z_b = -H + \tilde{h}(\xi, t)$ and the rigid lid at $z = 0$. We assume that the lateral boundaries are always far from the ridge where the water depth is H . Let

$$\bar{E} = \iint_{\mathcal{D}} E_{kp} + E_a \, d\xi \, dz \quad (3.16)$$

be the total perturbation energy. Taking the time derivative and making use of (3.14) and the boundary conditions $\tilde{\mathbf{u}} \cdot \hat{\mathbf{n}} = 0$ along the upper boundary and $\mathbf{u} \cdot \hat{\mathbf{n}} \, ds = -\bar{U}_b h' \, dx$ along the lower boundary, leads to

$$\frac{d\bar{E}}{dt} = (KE_f + APE_f + W) \Big|_{\xi_r}^{\xi_l} + \bar{U}_b(t) \int_{\xi_l}^{\xi_r} p_b h'(x(\xi, t)) \, d\xi, \quad (3.17)$$

where $p_b(\xi, t) = p_d(\xi, -H + \tilde{h}(\xi, t), t)$ is the pressure perturbation evaluated along the bottom. Here

$$\left. \begin{aligned} KE_f &= \int_{-H}^0 u E_k \, dz, \\ APE_f &= \int_{-H}^0 u E_a \, dz \end{aligned} \right\} \quad (3.18)$$

are the vertically integrated kinetic and available potential energy flux densities, and

$$W = \int_{-H}^0 u p_d \, dz \quad (3.19)$$

is the rate work is done by the pressure perturbation. We will refer to this as the work term. The total energy flux through a horizontal location x is $E_f = KE_f + APE_f + W$. The last term on the right-hand side of (3.17) is the generation, or conversion, term G (Khaliwala 2003; Kelly, Nash & Kunze 2010), which after integrating by parts (assuming $h(\xi_l) = h(\xi_r) = 0$) can be written as

$$G = \bar{U}_b(t) \int_{\xi_l}^{\xi_r} p_b h'(x(\xi, t)) \, d\xi = -\bar{U}_b(t) \int_{\xi_l}^{\xi_r} (p_\xi + p_z h') \Big|_{z_b} h(x(\xi, t)) \, d\xi. \quad (3.20)$$

The latter form is useful because the numerical model solves for the pressure gradients, not the pressure. Equation (3.17) states that the rate of change of energy in the domain \mathcal{D} is balanced by the flux of energy through the lateral boundaries plus the rate energy is injected into the system through the generation term which in the moving reference frame is the work done moving the topography.

In linear theory in the absence of a background current (3.17) would simplify to

$$\frac{d\bar{E}}{dt} = W \Big|_{\xi_l}^{\xi_r} + \bar{U}_b(t) \int_{\xi_l}^{\xi_r} p_d h'(x(\xi, t)) d\xi \quad (3.21)$$

because E_k and E_a are second-order, and p_d is first-order, in the perturbations. The energy fluxes KE_f and APE_f , which would then be third-order in amplitude, do not appear in linear theory. This is not the case when there is a background current. As the only flux term that appears in (3.21) is W , this term is often called the energy flux (Llewellyn Smith & Young 2002; Khaliwala 2003; Nash, Alford & Kunze 2005).

Using $u = \bar{U}(z) + u'$ we can split KE_f into terms of order 0–3 in the perturbation velocity fields via

$$\begin{aligned} KE_f = & \frac{1}{2} \int_{-H}^0 \bar{U}^3 dz + \frac{3}{2} \int_{-H}^0 \bar{U}^2 u' dz \\ & + \frac{1}{2} \int_{-H}^0 \bar{U} (3u'^2 + w^2) dz + \frac{1}{2} \int_{-H}^0 u' (u'^2 + w^2) dz. \end{aligned} \quad (3.22)$$

The first term can be ignored as it is independent of x and t hence the values at ξ_l and ξ_r cancel. Henceforth KE_f excludes this term. We split the remaining contributions into parts that are first-, second- and third-order in the perturbation velocities:

$$\left. \begin{aligned} KE_{f1} &= \frac{3}{2} \int_{-H}^0 \bar{U}^2 u' dz, \\ KE_{f2} &= \int_{-H}^0 \bar{U} \left(\frac{3}{2} u'^2 + \frac{1}{2} w^2 \right) dz, \\ KE_{f3} &= \frac{1}{2} \int_{-H}^0 u' (u'^2 + w^2) dz. \end{aligned} \right\} \quad (3.23)$$

In our simulations the third-order term and the contribution from w^2 to the second-order term are negligible.

The work term W can be split into two terms via

$$W = \int_{-H}^0 \bar{U} p_d dz + \int_{-H}^0 u' p_d dz = W_1 + W_2, \quad (3.24)$$

with W_1 and W_2 being of first and second order in the perturbation fields p_d and u' . Both make leading-order contributions to the tidally averaged work term because of the generation of second-order mean pressure fields.

Linear hydrostatic theory for linear stratifications in the absence of a background current predicts that the energy fluxes scale with

$$F_0 = \frac{1}{2} N U_0^2 h_0^2, \quad (3.25)$$

along with other factors depending on, for example, the ridge shape (Llewellyn Smith & Young 2002; Garrett & Kunze 2007), hence we will non-dimensionalize our energy fluxes by scaling them by F_0 .

4. Linear theory

In this section we develop a linear theory to predict the waves that are generated by periodic tidal flow over an isolated bathymetric feature in the presence of a steady surface trapped background current restricted to lie above the bathymetry so that waves are generated solely by tide–topography interactions. As mentioned above rotational effects are not included. The stratification and current are assumed to be stable (Richardson number greater than $1/4$) but are otherwise arbitrary. We follow the procedure used by Khatiwala (2003) and assume that the tidal flow consists of a single tidal constituent and seek periodic solutions. Multiple constituents can be considered via linear superposition. One difference between the fully nonlinear numerical simulations and the linear theory is that the former is an initial value problem while the latter is a periodic solution which can be viewed as the wave field in the limit $t \rightarrow \infty$.

As in the previous section we use a reference frame moving with the far field tidal current $\bar{U}_b(t)$. We assume small amplitude topography and waves and linearize about a steady background horizontal current $\bar{U}(z)$. This leads to the well-known equation for the vertical velocity

$$\left(\frac{\partial}{\partial t} + \bar{U} \frac{\partial}{\partial \xi}\right)^2 \nabla^2 w - \frac{d^2 \bar{U}}{dz^2} \left(\frac{\partial}{\partial t} + \bar{U} \frac{\partial}{\partial \xi}\right) w_\xi + N^2(z) w_{\xi\xi} = 0, \quad (4.1)$$

where the buoyancy frequency N is given by

$$N^2(z) = -g \frac{d\bar{\rho}}{dz}, \quad (4.2)$$

which for the theory is an arbitrary non-negative function of z . The linearized bottom boundary condition is

$$w = \bar{U}_b(t) \tilde{h}_\xi = \bar{U}_b(t) h'(x(\xi, t)) \quad \text{at } z = -H. \quad (4.3)$$

Searching for solutions of the form

$$w = e^{i(k\xi - \omega t)} \phi(z) \quad (4.4)$$

we obtain the Taylor–Goldstein equation for the vertical structure $\phi(z)$,

$$\phi'' + \left(\frac{N^2(z)}{(\bar{U}(z) - c)^2} - \frac{\bar{U}''(z)}{\bar{U}(z) - c} - k^2 \right) \phi = 0, \quad (4.5)$$

where $c = \omega/k$ and primes denote differentiation with respect to z . Solutions satisfying the homogeneous boundary conditions $\phi(-H) = \phi(0) = 0$ consist of two types: the discrete spectrum, or eigenmodes, $\{\phi_n^\pm, c_n^\pm\}$ with phase speeds $c_n^+ > \max\{\bar{U}\}$ or $c_n^- < \min\{\bar{U}\}$; and the continuous spectrum with $c \in [\min\{\bar{U}\}, \max\{\bar{U}\}]$. Figure 4 is a schematic showing the discrete and continuous spectrums when $U_{\min} < 0 < U_{\max}$.

We use the Fourier transform defined as

$$\hat{f}(k) = \int_{-\infty}^{\infty} f(\xi) e^{-ik\xi} d\xi \quad (4.6)$$

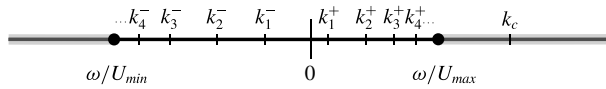


FIGURE 4. Spectrum of the Taylor–Goldstein equation for frequency $\omega > 0$. Shown are the first four positive and negative eigenvalues k_n^+ and k_n^- (the discrete spectrum) with limit points at ω/U_{max} and ω/U_{min} respectively. Also shown is point in the continuous spectrum $k_c = \omega/U(z)$ for a value of z for which $0 < U(z) < U_{max}$. The grey bars indicate regions where the continuous spectrum lies: as z varies across the water depth k_c sweeps out the grey regions $k < \omega/U_{min}$ and $k > \omega/U_{max}$. As $U_{min} \rightarrow 0$ the limit point at $\omega/U_{min} \rightarrow -\infty$.

to transform the bottom boundary condition (4.3), yielding (Bell 1975a; Khatiwala 2003)

$$\hat{w}(k, -H, t) = -\hat{h}(k) \sum_{n=-\infty}^{\infty} in\omega_0 J_n \left(\frac{-kU_0}{\omega_0} \right) e^{-in\omega_0 t}, \quad (4.7)$$

where the tidal current is

$$\bar{U}_b(t) = U_0 e^{i\omega_0 t} \quad (4.8)$$

and J_n is the Bessel function of the first kind of order n . Based on this we look for a series solution for $\hat{w}(k, z, t)$ of the form

$$\hat{w}(k, z, t) = - \sum_{n=-\infty}^{\infty} in\omega_0 \hat{h}(k) J_n \left(\frac{-kU_0}{\omega_0} \right) W_n(k, z) e^{-in\omega_0 t}, \quad (4.9)$$

where the functions $W_n(k, z)$ are solutions of the Taylor–Goldstein equation for waves of frequency $n\omega_0$,

$$\mathcal{L}_n[W_n] \equiv W_n'' + \left(\frac{k^2 N^2(z)}{(n\omega_0 - k\bar{U}(z))^2} + k \frac{\bar{U}''(z)}{(n\omega_0 - k\bar{U}(z))} - k^2 \right) W_n = 0, \quad (4.10)$$

satisfying the boundary conditions

$$\left. \begin{aligned} W_n(k, 0) &= 0, \\ W_n(k, -H) &= 1. \end{aligned} \right\} \quad (4.11)$$

So far the derivation has followed Khatiwala (2003) who considered constant N and did not include background currents. Under these conditions the Taylor–Goldstein equation can be solved analytically. With a sheared background current or non-constant $N(z)$ equations (4.10)–(4.11) must be solved numerically. To do so we numerically solve the initial value problem

$$\left. \begin{aligned} \mathcal{L}_n[\phi_n] &= 0, \\ \phi_n(k, 0) &= 0, \\ \phi_n'(k, 0) &= 1 \end{aligned} \right\} \quad (4.12)$$

to get $\phi_n(k, z)$ on $z \in [-H, 0]$ and then let

$$W_n(k, z) = \frac{\phi_n(k, z)}{\phi_n(k, -H)}. \quad (4.13)$$

Here k varies continuously over the real line. The values of k for which $\phi_n(k, -H) = 0$ give the eigenmodes, for which W_n is singular.

Taking the inverse Fourier transform of (4.9) we obtain

$$w(\xi, z, t) = \sum_{\substack{-\infty \\ n \neq 0}}^{\infty} \frac{-in\omega_0}{2\pi} e^{-in\omega_0 t} \int_{-\infty}^{\infty} \hat{h}(k) J_n \left(\frac{-kU_0}{\omega_0} \right) \frac{\phi_n(k, z)}{\phi_n(k, -H)} e^{ik\xi} dk. \quad (4.14)$$

The contributions to the integral are of two types: (i) contributions from the poles where $\phi_n(k, -H) = 0$ and (ii) possibly from the continuous spectrum. Without loss of generality we will assume that $U_{\min} = \min\{\bar{U}\} \leq 0$ and $U_{\max} = \max\{\bar{U}\} > 0$. In this paper we do not consider the continuous spectrum as the wave field appears to be well predicted by the discrete spectrum alone.

4.1. Contribution from the discrete spectrum

For $n\omega_0/U_{\min} < k < n\omega_0/U_{\max}$, or, if $U_{\min} = 0$ as is the case in our numerical simulations, for $-\infty < k < n\omega_0/U_{\max}$, the contributions to the integral from the discrete spectrum come from the poles of the integrand, i.e. the values of k for which $\phi_n(k, -H) = 0$. The corresponding ϕ_n are the eigenmodes. Let k_{nm}^+ , $m = 1, 2, \dots$ be the positive zeros of $\phi_n(k, -H)$ and let k_{nm}^- , $m = 1, 2, \dots$ be the negative zeros. Ordering them in decreasing phase speed magnitude we have

$$k_{n1}^+ < k_{n2}^+ < k_{n3}^+ < \dots \quad \text{with } k_{nm}^+ \rightarrow \frac{n\omega_0}{U_{\max}} \text{ as } m \rightarrow \infty \quad (4.15)$$

and

$$-k_{n1}^- < -k_{n2}^- < -k_{n3}^- < \dots \quad \text{with } -k_{nm}^- \rightarrow -\frac{n\omega_0}{U_{\min}} \text{ as } m \rightarrow \infty. \quad (4.16)$$

In our examples with $U_{\min} = 0$ we have $k_{nm}^- \rightarrow -\infty$ as $m \rightarrow \infty$. In the above n refers to the wave harmonic, i.e. waves of frequency $n\omega_0$, while m is the eigenmode index. For positive/negative n , the wavenumbers k_{nm}^+ correspond to rightward/leftward propagating waves. This is reversed for wavenumbers k_{nm}^- .

If $\{\phi_n(k, z), k\}$ is the solution of the Taylor–Goldstein equation subject to boundary conditions (4.12) for $\omega = n\omega_0$ then $\{\phi_n(k, z), -k\}$ is a solution for $\omega = -n\omega_0$ since the problem for ϕ is invariant under concurrent changes of the signs of ω and k . Thus

$$k_{nm}^+ = -k_{-nm}^- \quad (4.17)$$

and

$$\phi_n(k, z) = \phi_{-n}(-k, z). \quad (4.18)$$

Differentiating the latter gives

$$\frac{\partial \phi_n}{\partial k}(k, z) = -\frac{\partial \phi_{-n}}{\partial k}(-k, z). \quad (4.19)$$

Rightward propagating waves. Contributions to rightward propagating waves come from the positive eigenvalues for frequencies $n\omega_0 > 0$ and from the negative eigenvalues for frequencies $-n\omega_0 < 0$ (Khaliwala 2003). Use of the residue theorem

and the symmetries given above leads to the following contributions to u and w from the discrete spectrum:

$$w_{ds}(\xi, z, t) = 2 \sum_{n=1}^{n_0} \sum_{m=1}^{\infty} \left\{ n\omega_0 J_n \left(\frac{-k_{nm}^+ U_0}{\omega_0} \right) \frac{\phi_n(k_{nm}^+, z)}{\frac{\partial \phi_n}{\partial k}(k_{nm}^+, -H)} \operatorname{Re} \{ \hat{h}(k_{nm}^+) e^{i(k_{nm}^+ \xi - n\omega_0 t)} \} \right\}, \quad (4.20)$$

$$u_{ds}(\xi, z, t) = 2 \sum_{n=1}^{n_0} \sum_{m=1}^{\infty} \left\{ n\omega_0 J_n \left(\frac{-k_{nm}^+ U_0}{\omega_0} \right) \frac{\phi'_n(k_{nm}^+, z)}{\frac{\partial \phi_n}{\partial k}(k_{nm}^+, -H)} \operatorname{Re} \left\{ i \frac{\hat{h}(k_{nm}^+)}{k_{nm}^+} e^{i(k_{nm}^+ \xi - n\omega_0 t)} \right\} \right\} \quad (4.21)$$

for $\xi > 0$. Here ϕ'_n denotes differentiation with respect to z .

Leftward propagating waves. For the leftward propagating waves we have

$$w_{ds}(\xi, z, t) = -2 \sum_{n=1}^{n_0} \sum_{m=1}^{\infty} \left\{ n\omega_0 J_n \left(\frac{-k_{nm}^- U_0}{\omega_0} \right) \frac{\phi_n(k_{nm}^-, z)}{\frac{\partial \phi_n}{\partial k}(k_{nm}^-, -H)} \operatorname{Re} \{ \hat{h}(k_{nm}^-) e^{i(k_{nm}^- \xi - n\omega_0 t)} \} \right\}, \quad (4.22)$$

$$u_{ds}(\xi, z, t) = -2 \sum_{n=1}^{n_0} \sum_{m=1}^{\infty} \left\{ n\omega_0 J_n \left(\frac{-k_{nm}^- U_0}{\omega_0} \right) \frac{\phi'_n(k_{nm}^-, z)}{\frac{\partial \phi_n}{\partial k}(k_{nm}^-, -H)} \operatorname{Re} \left\{ i \frac{\hat{h}(k_{nm}^-)}{k_{nm}^-} e^{i(k_{nm}^- \xi - n\omega_0 t)} \right\} \right\} \quad (4.23)$$

for $\xi < 0$.

The sum over n has an upper limit because eigenmodes only exist for a finite number of frequencies (e.g. $n\omega_0 < N$ for N constant and $\tilde{U}(z) = 0$).

To cast the results into non-dimensional terms, because the eigenmodes are scaled so that $\phi'_n(k, 0) = 1$, we take ϕ'_n to be dimensionless. Denoting dimensionless variables with a tilde we have

$$(\phi_n, \hat{h}, k, \omega_0, U_0) = \left(H\tilde{\phi}_n, H^2\tilde{\hat{h}}, \frac{\tilde{k}}{H}, N\tilde{\omega}_0, NH\tilde{U}_0 \right). \quad (4.24)$$

Figure 5 compares phase speeds for the first 20 modes for three different background currents (the maximum difference between the phase speeds and corresponding group velocities is less than 2%). For the rightward propagating waves the phase speeds increase as current strength U_s or depth $-z_s$ is increased, decreasing monotonically to U_s as the mode number $m \rightarrow \infty$. The phase speeds of the upstream propagating waves decrease as the current is strengthened. They are much less sensitive to the strength or depth of the current. The phase speeds are not sensitive to the shear layer thickness d_s . Sample eigenmodes are shown in figure 6. For the leftward propagating modes $k < 0$ and $\omega_0 - \tilde{U}k$ is larger in the current than beneath it, hence the higher modes oscillate more rapidly beneath the current. For the rightward propagating modes $k > 0$ and $\omega_0 - \tilde{U}k$ is smaller in the current. The higher modes now oscillate more rapidly in the current.

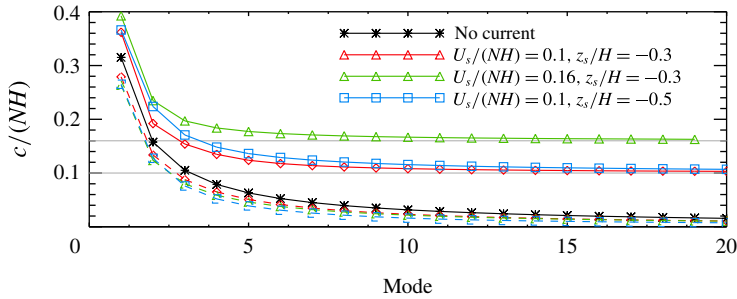


FIGURE 5. Wave phase speeds for waves of tidal frequency for various background currents. Horizontal lines indicate limiting values of 0.1 and 0.16. Solid curves are phase speeds of rightward propagating waves. Dashed are phase speeds (absolute value) of leftward propagating waves.

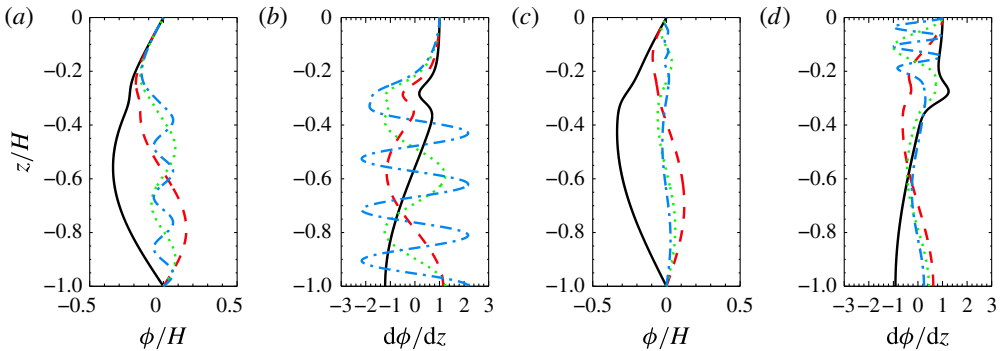


FIGURE 6. Eigenmodes (a,c) 1, 2, 4 and 8 and their derivatives (b,d) for waves of tidal frequency for background current used in sets 1–3 for which $(U_s/(NH), z_s/H, d_s/H) = (0.1, -0.3, -0.06)$. (a,b) Leftward propagating waves. (c,d) Rightward propagating waves.

The presence of the current has a significant impact on the modal wavenumbers which in turn affects the relative contributions of the different modes. Contributions to w_{ds} come from $\hat{h}(k)$ which depends only on the properties of the bathymetry, $J_n(-kU_0/\omega_0)$ which depends on the tidal current amplitude U_0 and frequency ω_0 , and on the eigenmodes which are determined by the stratification and background currents. We now briefly consider the contributions from some of the factors in expressions (4.20)–(4.23).

The Gaussian ridge (2.2) has Fourier transform

$$\hat{h}(k) = \sqrt{\pi}ah_0e^{-(a^2k^2)/4}. \quad (4.25)$$

Figure 7(a) shows $f_1 = \hat{h}(k)/\hat{h}(0)$ for $a/H = 1.2, 1.8, 2.4, 3.0$ and 3.6 . As a/H increases the spectrum gets narrower. The values for the discrete spectrum for waves of tidal frequency from set 4 (see table 1) are indicated with red stars. Rightward propagating waves ($k > 0$) have limiting wavenumber $k_{lim}H = (\omega_0/U_{max})H \approx 1.4$. For $a/H = 1.2$ and 1.8 all eigenmodes make a significant contribution. For the widest topography only the first 3 eigenmodes do. For the leftward propagating waves $k_{lm}^- \rightarrow -\infty$ as $m \rightarrow \infty$ and fewer eigenmodes make a significant contribution.

These values are multiplied by $J_1(-(kU_0/\omega_0))$. In our case $U_0 = U_s/10$ so this factor is equal to $J_1(-(k/10k_{lim})) \approx k/20k_{lim}$. Figure 7(b) shows $f_2 = \hat{h}(k)J_1(-(kU_0/\omega_0))/\hat{h}(0)$. The Bessel function makes contributions from higher harmonics negligible.

Including the contribution to the vertical velocity from $\partial\phi_n/\partial k$ (figure 7c) we see that for narrow ridges several modes make higher contributions than mode-one does. The highest contributions to the horizontal velocity, which has an extra division by k , are more skewed to the low modes. For $a/H = 1.2$ and 1.8 the largest contributions for the rightward propagating waves are from modes 2, 3 and 1 (figure 7d). For wider ridges mode-one dominates. Trends are similar for the leftward propagating waves. The signs of the $\partial\phi_n/\partial k$ terms alternate with mode number (we use absolute values in the figure) and the actual relative contributions will depend on the relative phases of the waves.

4.2. First-order linear energy flux

Formally, a small amplitude parameter ϵ denoting the ridge amplitude divided by the water depth can be introduced. The generated waves depend on ϵ . Expanding the perturbation velocity u in a perturbation expansion in powers of ϵ we have

$$u = u(\xi, z, t; \epsilon) = \epsilon u^{(0)}(\xi, z, t) + \epsilon^2 u^{(1)}(\xi, z, t) + \dots, \quad (4.26)$$

with similar expansions for w , p and ρ . The above linear solution gives the $O(\epsilon)$ terms. These terms are by assumption periodic in t , being comprised of terms with frequencies $n\omega_0 < N$. In particular their tidal average is zero. Nonlinear interactions produce $O(\epsilon^2)$ terms with non-zero tidal averages: mean second-order horizontal currents and second-order mean density and pressure perturbations.

In terms of this perturbation expansion the work term, away from the ridge, is

$$W = \int_{-H}^0 u p_d dz = \epsilon \int_{-H}^0 \bar{U}(z) p^{(0)} dz + \epsilon^2 \int_{-H}^0 (\bar{U}(z) p^{(1)} + u^{(0)} p^{(0)}) dz + O(\epsilon^3). \quad (4.27)$$

Using $\langle \cdot \rangle$ to denote tidally averaged values, we have

$$\langle W \rangle = \epsilon^2 \int_{-H}^0 (\bar{U}(z) \langle p^{(1)} \rangle + \langle u^{(0)} p^{(0)} \rangle) dz + O(\epsilon^3) \quad (4.28)$$

since the mean of $p^{(0)}$ is zero. Similarly, the tidally averaged kinetic energy flux is

$$\langle KE_f \rangle = \epsilon^2 \int_{-H}^0 \left(\frac{3}{2} \bar{U}(z) \langle u^{(1)} \rangle + \frac{1}{2} \bar{U}(z) \langle 3(u^{(0)})^2 + (w^{(0)})^2 \rangle \right) dz + O(\epsilon^3), \quad (4.29)$$

and the available potential energy flux is

$$\langle APE_f \rangle = \int_{-H}^0 \bar{U}(z) \langle E_a \rangle dz = \epsilon^2 \int_{-H}^0 \frac{g^2}{2N^2} \bar{U}(z) \langle (\rho^{(0)})^2 \rangle dz + O(\epsilon^3). \quad (4.30)$$

For $\bar{U}(z) \neq 0$ the leading-order mean kinetic and available potential energy fluxes are of the same order as the work term and the mean second-order pressure and horizontal currents contribute to the leading-order terms.

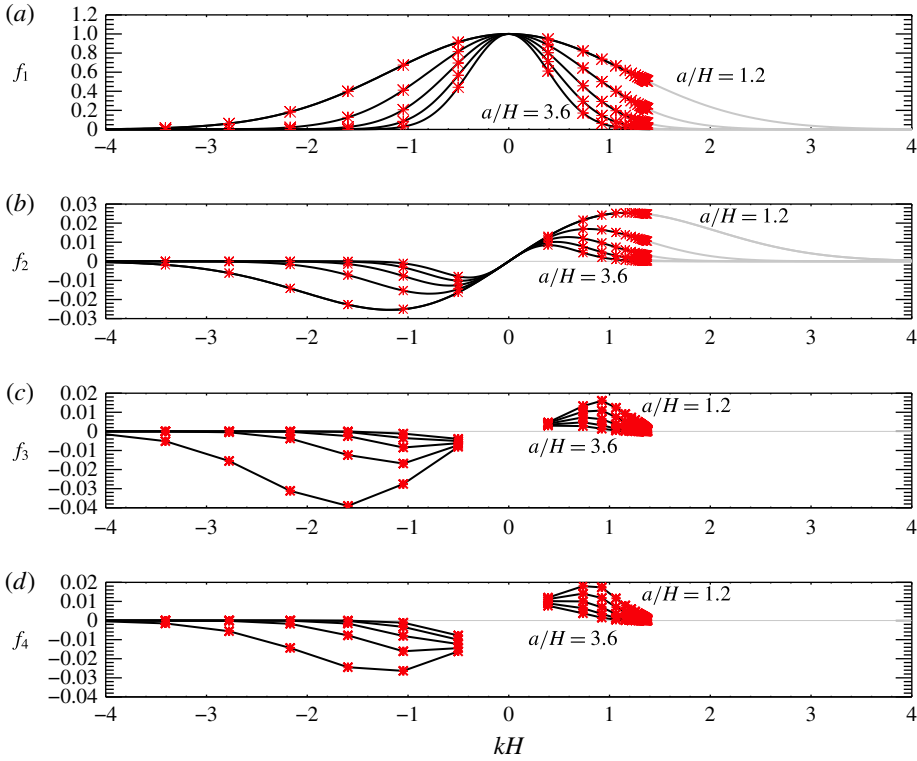


FIGURE 7. Partial coefficients in expressions for theoretical values for u_{ds} and w_{ds} for waves of tidal frequency ($n = 1$). (a) $f_1 = \hat{h}(k)/\hat{h}(0)$ for the Gaussian ridge. Curves are for $a/H = 1.2, 1.8, 2.4, 3.0$ and 3.6 . Red stars show the discrete spectrum for background current (2.4) for $U_s/(NH) = 0.1$, $d_s/H = 0.08$ and $z_s/H = -0.3$. (b) $f_2 = (\hat{h}(k)/\hat{h}(0))J_1(-kU_0/\omega_0)$. (c) $f_3 = (\hat{h}(k)/\hat{h}(0))J_1(-kU_0/\omega_0)/|\partial\phi_1/\partial k|$. (d) $f_4 = (\hat{h}(k)/\hat{h}(0))J_1(-kU_0/\omega_0)/|\partial\phi_1/\partial k|k$. In panels (c,d) values shown are for k_{1m}^\pm . In panels (a,b) black curves show range of k values to which the discrete spectrum belong; grey shows the range of the continuous spectrum.

Because the second and higher harmonics are very small in our simulations in the following we assume all first-order terms are of tidal frequency. From linear theory we can calculate

$$\langle W \rangle_{lin} = \int_{-H}^0 \langle u^{(0)} p^{(0)} \rangle dz = \sum_{i=1}^M \sum_{j=1}^M \int_{-H}^0 \langle u_i p_j \rangle dz, \quad (4.31)$$

where M is the number of modes included and u_i and p_i are the contributions of mode- i to $u^{(0)}$ and $p^{(0)}$. Let

$$W_{ij}^l = \int_{-H}^0 \langle u_i p_j \rangle dz, \quad (4.32)$$

with the superscript l denoting linear theory. From (4.21) and (4.23), using the fact that \hat{h} is real for our symmetric Gaussian ridge, let

$$u_i = A_i \sin(\theta_i(x, t)) \phi_i'(z), \quad (4.33)$$

where $\theta_i(x, t) = k_i x - \omega_0 t$ for all i since the modes are in phase at $x = t = 0$. Then incompressibility gives

$$w_i = -k_i A_i \cos(\theta_i) \phi_i(z) \quad (4.34)$$

and from the horizontal momentum equation

$$p_i = \frac{A_i}{k_i} \sin(\theta_i) [(c_i - \bar{U}) \phi'_i + \bar{U}'(z) \phi_i(z)] \quad (4.35)$$

hence

$$u_i p_j = \frac{A_i A_j}{2k_j} [\cos(\theta_i - \theta_j) - \cos(\theta_i + \theta_j)] [(c_j - \bar{U}) \phi'_j + \bar{U}'(z) \phi_j(z)] \phi'_i(z). \quad (4.36)$$

Taking the time average we obtain

$$W'_{ij} = \frac{A_i A_j}{2k_j} \cos((k_i - k_j)x) \int_{-H}^0 [(c_j - \bar{U}) \phi'_j + \bar{U}'(z) \phi_j(z)] \phi'_i(z) dz. \quad (4.37)$$

In the absence of a background current $W'_{ij} = 0$ for $i \neq j$ because the eigenmodes ϕ'_i are orthogonal for waves of a fixed frequency but in general this is not the case. Hence the cross-terms W'_{ij} for $i \neq j$ make a spatially varying contribution to the energy flux.

Similarly, the leading-order contribution to the kinetic energy flux from the leading-order velocity perturbation is

$$\begin{aligned} \langle KE_f \rangle_{lin} &= \int_{-H}^0 \frac{1}{2} \bar{U} \langle 3(u^{(0)})^2 + (w^{(0)})^2 \rangle dz \\ &= \sum_{i,j} \frac{3}{4} A_i A_j \cos((k_i - k_j)x) \int_{-H}^0 \bar{U}(z) \phi'_i \phi'_j dz \\ &\quad + \sum_{i,j} \frac{1}{4} A_i A_j k_i k_j \cos((k_i - k_j)x) \int_{-H}^0 \bar{U}(z) \phi_i \phi_j dz. \end{aligned} \quad (4.38)$$

The second term, which is the flux of vertical kinetic energy, is negligible in our simulations. Let

$$KE'_{fij} = \frac{1}{4} A_i A_j \cos((k_i - k_j)x) \int_{-H}^0 \bar{U}(z) [3\phi'_i \phi'_j + k_i k_j \phi_i \phi_j] dz. \quad (4.39)$$

As for the work term, the cross-terms with $i \neq j$ are generally non-zero. The linearized density equation gives

$$\rho_i = \frac{A_i}{g} \sin \theta_i \frac{N^2}{c_i - \bar{U}} \phi_i, \quad (4.40)$$

the leading-order contribution to the available potential energy flux is

$$\langle APE_f \rangle_{lin} = \sum_{i,j} APE'_{fij}, \quad (4.41)$$

where

$$APE'_{fij} = \frac{1}{4} A_i A_j \cos((k_i - k_j)x) \int_{-H}^0 \frac{N^2}{(c_i - \bar{U})(c_j - \bar{U})} \phi_i \phi_j dz. \quad (4.42)$$

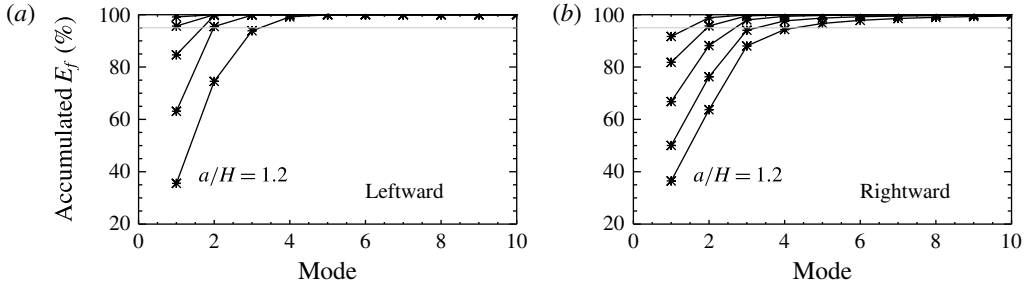


FIGURE 8. Accumulated percentage of total linear energy flux $100 \times \sum_{j=1}^{n_m} E_{fj}^l / \sum_{j=1}^{16} E_{fj}^l$ versus number of modes n_m . (a) Leftward propagating waves. (b) Rightward propagating waves. Results are shown for $a/H = 1.2, 1.8, 2.4, 3.0$ and 3.6 . As a/H increases the fraction of the total energy flux in the first mode increases.

Figure 8 shows the accumulated percentage of the total linear energy flux contributed by the squared terms as a function of the number of modes, i.e. $\sum_{i=1}^M E_{fi}^l$ where $E_{fi}^l = W_{ii}^l + KE_{fii}^l + APE_{fii}^l$. Only the squared terms with $i=j$ are included because the cross-terms vary with x with average values of zero. More modes are needed to estimate the total energy flux for the downstream (rightward) propagating waves than for the upstream (leftward) propagating waves and the narrower the ridge the more modes are needed. As we will see below, the linear theory does not accurately predict the energy flux for narrow ridges.

5. Validation tests

Validation tests were done by checking convergence and how well the energy balance equation (3.17) is satisfied. Here we present results for the ridge with $a/H = 1.8$ from set 4. The minimum Richardson number is 1.8 and the slope is sub-critical with $\gamma = 0.34$.

Simulations done with horizontal resolutions of $\Delta x/H = 0.02$ and 0.01 and time steps of $N\Delta t = 0.069$ and 0.0345 showed that $\Delta x/H = 0.02$ and $N\Delta t = 0.069$ were generally adequate. For the strongest current we use $N\Delta t = 0.0345$. A uniform vertical grid spacing with 250 grid cells was used, 400 grid cells giving similar results. This gives a vertical resolution of $\Delta z/H = 0.004$ in the deep water.

Figure 9 compares terms in the energy balance equation (3.17) for a Lagrangian region spanning the ridge. The dominant balance is between $d\bar{E}/dt$ and $W|_{\xi_l}^{\xi_r}$ (figure 9a). A major contribution to the energy \bar{E} is associated with the time varying barotropic currents over the ridge. In conjunction there are large oscillating pressure gradients accelerating the barotropic currents. These form the dominant balance in the energetics over the ridge. Panel (b) shows that $d\bar{E}/dt - W|_{\xi_l}^{\xi_r}$ is largely balanced by $KE_f|_{\xi_l}^{\xi_r}$. Panel (c) shows that $d\bar{E}/dt - W|_{\xi_l}^{\xi_r} - KE_f|_{\xi_l}^{\xi_r}$ and the conversion term G are virtually identical. Panel (d) shows the sum of all the terms $d\bar{E}/dt - W|_{\xi_l}^{\xi_r} - KE_f|_{\xi_l}^{\xi_r} - APE_f|_{\xi_l}^{\xi_r} - G$ which should theoretically be equal to zero. Peak values are approximately 0.016 compared with peak values of approximately 60 for $d\bar{E}/dt$ so the energy budget is satisfied to within approximately 0.025 %.

The energy balance in a Lagrangian domain to the right of the ridge was also analysed using flux lines with mean positions of 5.4 and 9.4 (not shown). The length of this region is 25 % of the wavelength of a rightward mode-one tidal frequency wave.

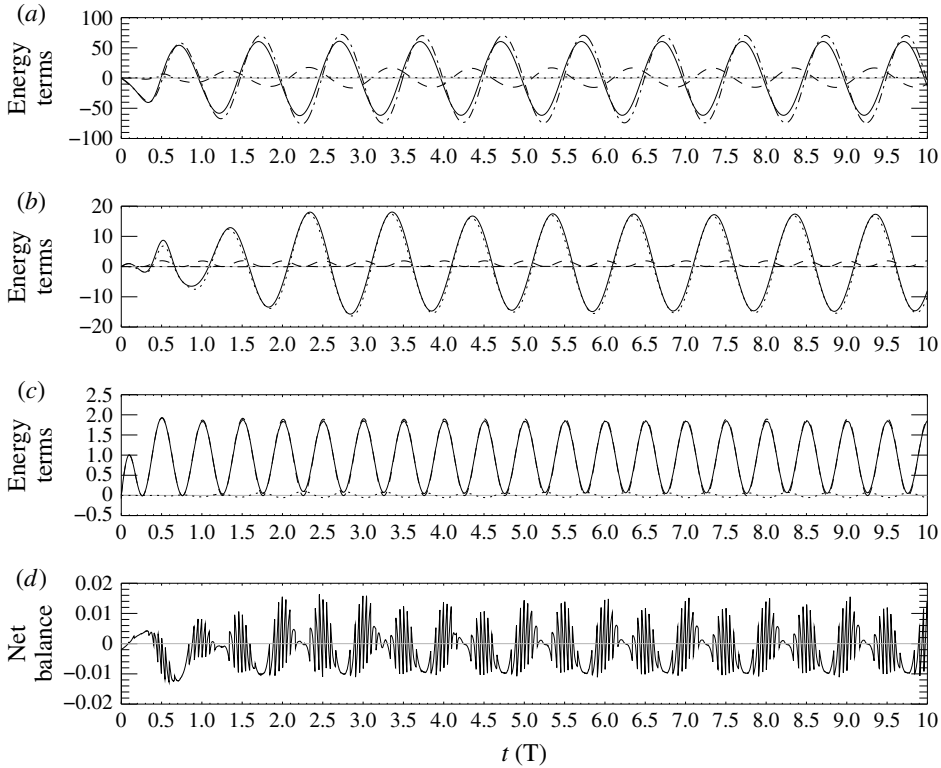


FIGURE 9. Energy balance in a Lagrangian domain spanning the ridge. Values non-dimensionalized by the flux scale F_0 . Time in tidal periods. The domain lies between flux lines with mean positions at $x/H = \pm 5.4$ for a case with $U_s/(NH) = 0.1$, $h_0/H = 0.1$ and $a/H = 1.8$. Resolution is $dx/H = 0.02$ in the horizontal with $J = 250$. Time step $N\Delta t = 0.069$. (a) $d\bar{E}/dt$ (solid), $W|_{\xi_r}^{\xi_l}$ (dash-dot), $KE_f|_{\xi_r}^{\xi_l}$ (dash), conversion term G (dots). (b) $d\bar{E}/dt - W|_{\xi_r}^{\xi_l}$ (solid), $KE_f|_{\xi_r}^{\xi_l}$ (dots), G (dashed), $APE_f|_{\xi_r}^{\xi_l}$ (dash-dots). The latter is indistinguishable from zero. (c) $d\bar{E}/dt - W|_{\xi_r}^{\xi_l} - KE_f|_{\xi_r}^{\xi_l}$ (solid black), G (dashed), $APE_f|_{\xi_r}^{\xi_l}$ (dots). The first two curves are barely distinguishable. (d) Net balance $d\bar{E}/dt - W|_{\xi_r}^{\xi_l} - G - KE_f|_{\xi_r}^{\xi_l} - APE_f|_{\xi_r}^{\xi_l}$.

Off the ridge the conversion term is zero and the energy fluctuations are not affected by flow constriction over the ridge. The dominant balance is between $d\bar{E}/dt$, $W|_{\xi_r}^{\xi_l}$ and $KE_f|_{\xi_r}^{\xi_l}$, with the latter two approximately half the size of $d\bar{E}/dt$. Peak values of the net balance $d\bar{E}/dt - W|_{\xi_r}^{\xi_l} - KE_f|_{\xi_r}^{\xi_l} - APE_f|_{\xi_r}^{\xi_l}$ are approximately 0.01, approximately 0.03 % of the peak values of $d\bar{E}/dt$ so the energy budget in this off-ridge domain is also satisfied to high accuracy.

Figure 10 compares tidally averaged values in the off-ridge Lagrangian domain. Panel (a) shows the work terms and kinetic energy fluxes at the left and right boundaries. All undergo oscillations with a period of about $4T$ as well as longer period fluctuations. The net energy balance, $\langle d\bar{E}/dt \rangle - \langle W \rangle|_{\xi_r}^{\xi_l} - \langle KE_f \rangle|_{\xi_r}^{\xi_l} - \langle APE_f \rangle|_{\xi_r}^{\xi_l}$, has a typical value of about 0.01 (panel c) and does not undergo significant oscillations. This is approximately 0.5 % of a typical value for $\langle W \rangle|_{\xi_l}$.

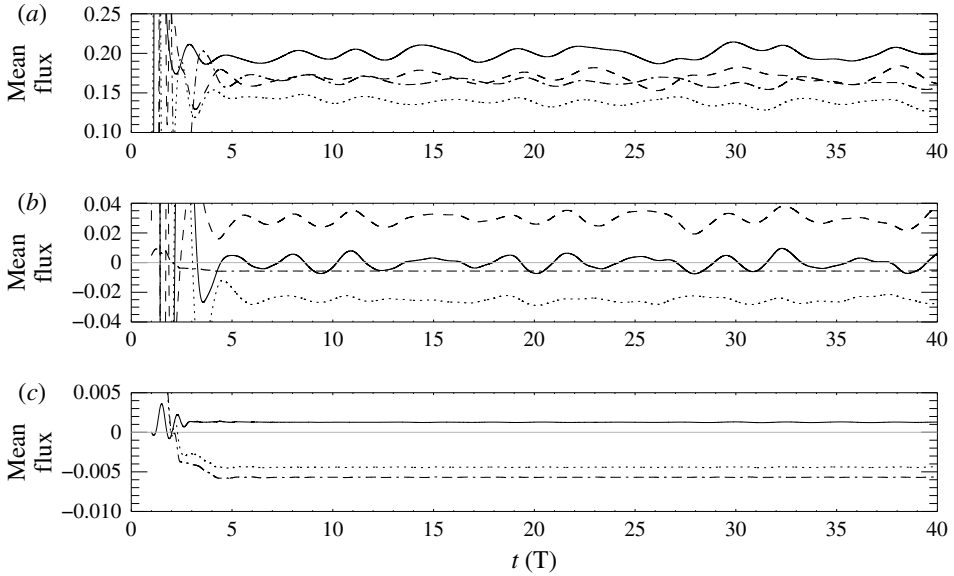


FIGURE 10. Energy balance in a Lagrangian domain off the ridge: tidal averages scaled by F_0 . The domain lies between flux lines with mean positions at $x/H = 5.4$ and 9.4 for the same case as in figure 9. (a) $\langle W|_{\xi_l} \rangle$ (solid), $\langle W|_{\xi_r} \rangle$ (dashed), $\langle KE_f|_{\xi_l} \rangle$ (dots) and $\langle KE_f|_{\xi_r} \rangle$ (dash-dot). (b) $\langle d\bar{E}/dt \rangle$ (solid black), $\langle W|_{\xi_r}^{\xi_l} \rangle$ (dashed), $\langle KE_f|_{\xi_r}^{\xi_l} \rangle$ (dots), $\langle APE_f|_{\xi_r}^{\xi_l} \rangle$ (dash-dot). (c) Net energy flux: $\langle d\bar{E}/dt \rangle - \langle W|_{\xi_r}^{\xi_l} \rangle - \langle KE_f|_{\xi_r}^{\xi_l} \rangle - \langle APE_f|_{\xi_r}^{\xi_l} \rangle$ (solid); $\langle d\bar{E}/dt \rangle - \langle W|_{\xi_r} \rangle - \langle KE_f|_{\xi_r} \rangle$ (dots), $\langle APE_f|_{\xi_r}^{\xi_l} \rangle$ (dash-dot).

6. Results

Total energy fluxes averaged over one tidal period have temporal fluctuations with amplitudes of up to 10% of their mean values. Averaging over four periods significantly reduced these variations so henceforth we use four period averages, i.e. we now define the averaging operator as

$$\langle \cdot \rangle(t) = \frac{1}{4T} \int_{t-4T}^t \cdot dt. \quad (6.1)$$

Most simulations were limited to twenty periods and we will generally report results at this time. Some sets were run for forty tidal periods.

Figure 11 shows the total energy flux $\langle E_f \rangle$ and the contributions from $\langle W \rangle$, $\langle KE_f \rangle$ and $\langle APE_f \rangle$, as a function of distance from the ridge for the cases from set 4 with $a/H = 1.8$ (a,b) and $a/H = 6$ (c,d). The former is the case depicted in figure 2. Values at $t = 20T$ (symbols) and $40T$ (grey curves) are shown. Note the much larger range of flux values used for the left side of the narrow ridge (panel a).

To the left of the narrower ridge the work term $\langle W \rangle$ and kinetic energy flux $\langle KE_f \rangle$, which have opposite signs, undergo large spatial fluctuations with a period $x/H \approx 11.3$ (figure 11a). Peak values of the magnitudes of $\langle KE_f \rangle$ and $\langle W \rangle$ occur at the locations where the beams reflect off the surface (see figures 1a and 2). The fluctuation period is close to $2\pi/H(k_2 - k_1) = 11.5$, the wavelength associated with the interaction of mode-one and mode-two waves (the mode-one wavelength is 12.5). The locations of the minimum and maximum values are independent of time (not shown) with values

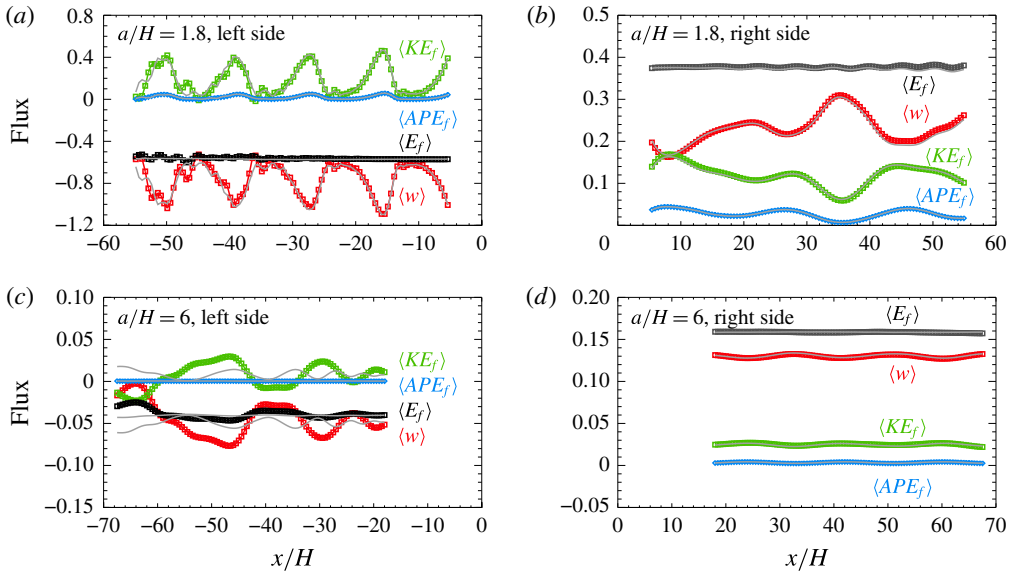


FIGURE 11. Tidally averaged energy fluxes at $t = 20T$ and $40T$ (averaged over previous four periods) scaled by F_0 as a function of distance for the cases from set 4 with $a/H = 1.8$ (a,b) and $a/H = 6$ (c,d). (a,c) Fluxes to left of ridge. (b,d) Fluxes to right of ridge. Note the much larger range of flux values in panel (a). Shown in each panel are the work term $\langle W \rangle$ (red), the total kinetic energy flux $\langle KE_f \rangle$ (green), the available potential energy flux $\langle APE_f \rangle$ (blue) and the total energy flux $\langle E_f \rangle$ (black). For comparison the grey curves show the fluxes at $t = 40T$.

undergoing relatively small temporal oscillations. To the left of the ridge the kinetic energy flux is positive, i.e. it is in the direction opposite to that of wave propagation. The total mean energy flux $\langle E_f \rangle$ is almost constant within a distance x/H of 40 from the ridge where the variations in $\langle W \rangle$ are almost completely balanced by variations in $\langle KE_f \rangle$. At larger distances from the ridge $\langle E_f \rangle$ has small spatial variations which grow with distance from the ridge. This appears to occur because the mean wave field beyond $x/H = -30$ is still in a transient state as indicated by the averaged energy flux after 40 tidal periods (grey curve). At $t = 20T$ linear theory predicts that the first 10 modes have reached $x/H = -30$ while only the first 5 have reached $x/H = -50$.

The fluxes to the right of the narrow ridge (figure 11b) undergo significantly smaller spatial variations on much longer length scales. Variations in $\langle W \rangle$ are significant and are again largely balanced by variations in $\langle KE_f \rangle$. Their pattern is very different to that on the left side of the ridge. $\langle KE_f \rangle$ is again in the direction of wave propagation. For both the leftward and rightward propagating waves it is in the direction of the background current as the dominant contribution is from the $\bar{U}u^2$ term (see below). The total energy flux at $t = 20T$ is almost constant out to $x/H = 54$ due to the larger downstream group velocities (note different scales used for fluxes to the left and right of the ridge: the oscillations far to the left of the ridge are much larger than those far to the right of the ridge). The different patterns in $\langle W \rangle$ and $\langle KE_f \rangle$ to the left and right of the ridge is connected to the different patterns in the mean wave field (see figure 2).

Figure 12 shows a breakdown of $\langle W \rangle$ and $\langle KE_f \rangle$ into their linear and quadratic parts (see (3.23) and (3.24)) for the narrower ridge ($a/H = 1.8$). The fluxes predicted

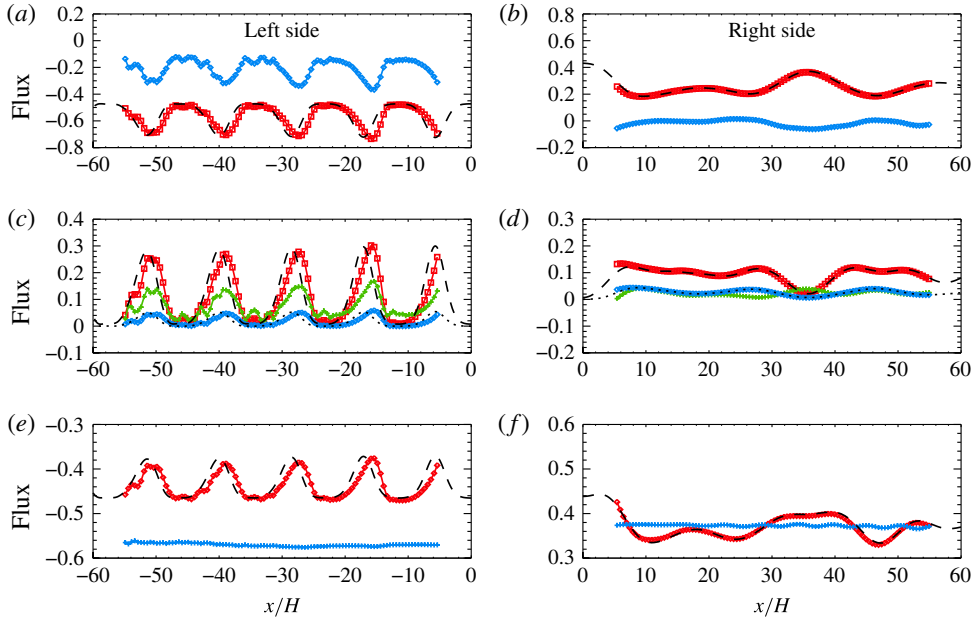


FIGURE 12. Comparison of theory and simulation results for $a/H = 1.8$ from set 4 (same case depicted in figure 11*a,b*). Shown are tidally averaged contributions to the energy fluxes at $t = 40T$ scaled by F_0 as a function of distance. Panels (a,c,e) show values to the left of the ridge. Panels (b,d,f) show values to the right of the ridge. (a,b) Contributions to the work term: $\langle W_1 \rangle$ (blue diamonds) and $\langle W_2 \rangle$ (red squares). The dashed black line shows $\langle W \rangle_{lin}$, the theoretical work term from first-order theory. (c,d) $\langle KE_{f1} \rangle$ (green), $\langle KE_{f2} \rangle$ (red) and $\langle APE_f \rangle$ (blue). The dashed and dotted lines show the theoretical kinetic energy flux $\langle KE_f \rangle_{lin}$ and available potential energy flux $\langle APE_f \rangle_{lin}$. (e,f) The total energy flux $\langle E_f \rangle$ from simulations (blue), contributions to the total energy flux that are second-order in the perturbation fields, $\langle W_2 + KE_{f2} + APE_f \rangle$ (red) and the total theoretical energy flux $\langle W + KE_f + APE_f \rangle_{lin}$ (black, dashed).

by first-order linear theory are also shown and they compare well with the second-order contributions. In particular, the linear first-order cross-terms, i.e. W_{ij}^l , KE_{fij}^l and APE_{fij}^l with $i \neq j$, account for the spatial variations in the fluxes in the numerical simulations. The fluxes that are linear in the perturbation fields, $\langle W_1 \rangle$ and $\langle KE_{f1} \rangle$, make important contributions to the total flux, i.e. the second-order mean wave fields are important. Panels (e,f) show the total energy flux, the total second-order flux $\langle W_2 + KE_{f2} + APE_f \rangle$, and the total energy flux from linear theory. The latter two are very similar. The second-order mean fields contribute almost 25 % of the total energy flux via the terms $\langle W_1 \rangle$ and $\langle KE_{f1} \rangle$ on the left side of the ridge. For $\langle KE_{f1} \rangle$ to be comparable to $\langle KE_{f2} \rangle$ we need $\langle u^{(1)} \rangle / NH$ to be comparable to $\langle (u^{(0)})^2 \rangle / \bar{U}_s / NH$. Using $\sqrt{\langle (u^{(0)})^2 \rangle} / NH \approx 0.004$ from figure 1, this gives $\langle u^{(1)} \rangle / NH \approx 1.6 \times 10^{-5}$ which is an order of magnitude less than the mean currents in the simulation, but given that the largest wave-induced mean currents are at the base of the shear layer where \bar{U} is smaller than U_s and downstream the mean currents alternate in sign, there is reasonable consistency.

Figure 11(c,d) shows the fluxes for the widest ridge ($a/H = 6$) from set 4. To the left of the ridge (panel c) the fluxes are approximately an order of magnitude smaller

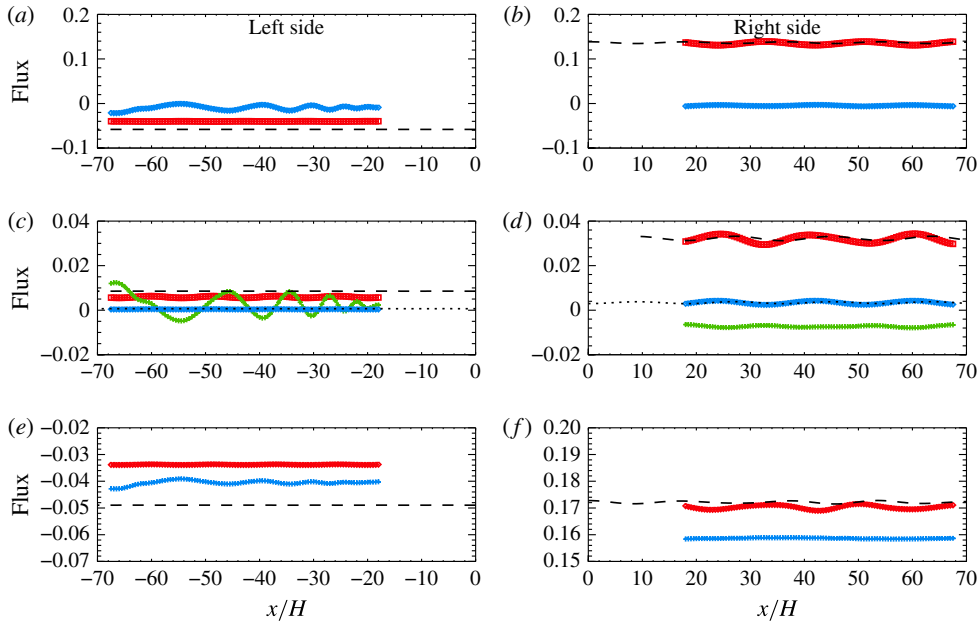


FIGURE 13. Same as figure 12 but for a wider ridge with $a/H = 6$.

than those for the narrower ridge. $\langle W \rangle$ and $\langle KE_f \rangle$ undergo oscillations which are large compared with their mean values. The fluxes at $t = 40T$ vary much less than those at $t = 20T$, in contrast to the narrower ridge, indicating that the flow takes longer to reach a steady state. The total flux near the ridge ($x/H = -18$) is approximately the same at $t = 20T$ and at $t = 40T$. To the right of the ridge the fluxes are much more uniform and do not change significantly between $20T$ and $40T$.

Figure 13 compares the first- and second-order contributions to the fluxes with the linear fluxes for the wide ridge. The spatial patterns in the fluxes are very different from the narrower ridge with much smaller spatial variations because energy is primarily in the first mode and the mean fields are much weaker (not shown). The differences between the second-order and theoretical fluxes (total and work term) is more significant for the wider ridge and is approximately halved for a smaller ridge with $h_0/H = 0.05$ (not shown).

Figure 14 shows the total energy flux $\langle E_f \rangle$ as a function of the ridge width for set 4. Fluxes in the absence of a background current are also shown. For narrow ridges the magnitudes of the upstream and downstream energy fluxes are increased and decreased by the current. For very wide ridges the opposite is the case though the effect of the current is much weaker. Fluxes to the left of the ridge increase in magnitude monotonically as the ridge gets narrower while the rightward fluxes have a maximum at $a/H \approx 2.7$. The ratio of the magnitudes of the rightward and leftward fluxes, R decreases monotonically from 4 for the widest ridge to 0.5 for the narrowest ridge.

Also shown in figure 14 are the energy fluxes predicted by first-order linear theory in which only the fluxes for individual modes have been used. That is, only the spatially uniform contribution

$$E_f^l = \sum_{i=1}^{n_{\max}} W_{ii}^l + K_{fi}^l + APE_{fi}^l \quad (6.2)$$

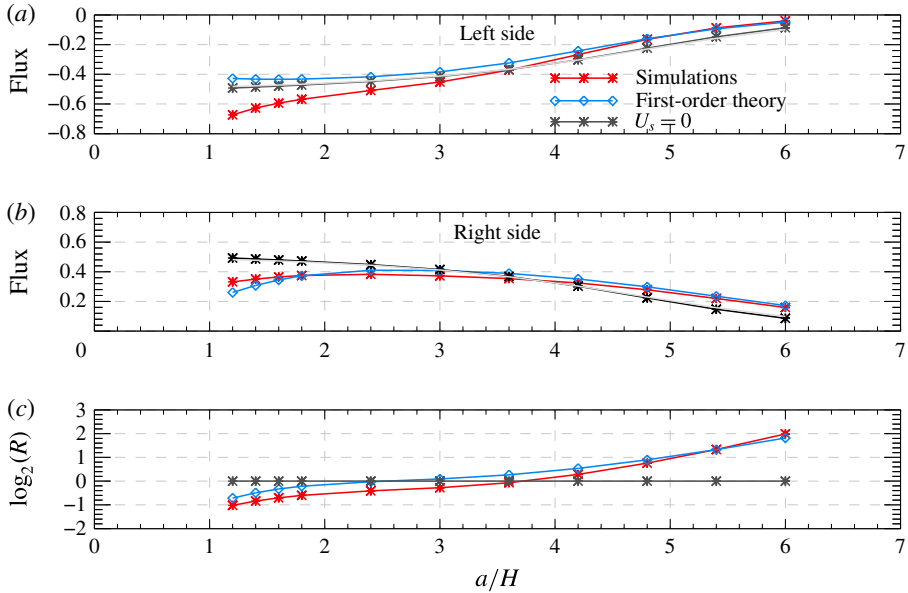


FIGURE 14. Tidally averaged energy fluxes at $t = 20T$ scaled by F_0 as a function of the ridge width a/H for set 4. Fluxes are calculated at: (a) $\xi = -3a/H$ for the leftward propagating waves; (b) $\xi = 3a/H$ for the rightward propagating waves. Values from the numerical simulations (red asterisks) and first-order linear theory (E_f^l , blue diamonds), excluding the spatially variable cross-terms, are shown. Also shown in (a,b) are the fluxes in the absence of a background current: simulation (black asterisks) and theory (grey curve). (c) Absolute value of the ratio of rightward to leftward fluxes (log base 2) from the numerical simulations from set 4 (green squares) and no background current (black asterisks).

is included. This truncated linear approximation predicts the energy flux for wide ridges ($a/H > \approx 4$) where only the lowest mode makes a significant contribution.

Figure 15 compares the total fluxes from the different sets of numerical simulations. Fluxes from sets 1–4 are shown in panels (a,d,g). Sets 1–3 differ only in the amplitude of the ridge and the results show that the fluxes scale closely with the amplitude squared. R increases with h_0 for the wide ridges (panel g). The results from sets 2 and 4, which differ only in the thickness of the shear layer, are nearly indistinguishable.

Figure 15(b,e,h) compare sets 4, 6 and 7 which differ primarily in the strength of the current. As expected, increasing the strength of the current increases the difference in the upstream and downstream fluxes. For narrow ridges the upstream fluxes are particularly sensitive to U_s , increasing rapidly as U_s is increased; $a/H \approx 3.6$ is the transition point below which the upstream flux exceeds the downstream flux.

Panels (c,f,i) compare the effects of increasing the depth of the current for two current strengths (sets 2, 5, 7 and 8). For the deeper currents (sets 5 and 8) the leftward fluxes are increased in magnitude for the narrow ridges and slightly reduced for the wider ridges. Because values are so small for the wide ridge this small reduction is large as a percentage: approximately 40% for $a/H = 6$ versus a 7% increase for $a/H = 1.2$. The downstream fluxes are reduced when the depth of the current is increased for ridges narrower than $a/H \approx 3$, increased for somewhat wider

ridges and unaffected when $a/H = 6$. As a consequence there is a large increase in R for wide ridges (panel *i*).

While the current introduces asymmetry in the upstream and downstream energy fluxes it does not significantly affect the total energy flux radiating away from the ridge (figure 16). This is not too surprising as the source of the internal waves at the bottom is well below the current. Indeed, the generation term is not sensitive to the presence of the currents. It is the projection of the perturbation onto the upstream and downstream propagating internal wave modes that is.

7. Summary

The effects of a surface trapped steady background current on internal wave generation by idealized inviscid tide–topography interactions has been investigated via two-dimensional nonlinear simulations. The uni-directional current $\bar{U}(z)$ has a simple structure having its maximum positive value at the surface, being almost vertically uniform in a surface layer before decaying monotonically to zero in a relatively thin shear layer. The minimum Richardson number of the background current is always well above 0.25. We also restrict attention to an idealized linear stratification for which there is a simple exact analytic expression for the available potential energy. Rotational effects are not included so that the problem is two-dimensional. Symmetric Gaussian ridges for a large range of widths were used. Most cases used a ridge amplitude equal to 10% of the water depth H the largest ridge being $0.2H$. All except for the tallest, narrowest ridge had subcritical slopes.

The background current results in asymmetries in the upstream (leftward) and downstream (rightward) wave fields and energy fluxes. The results did not show strong sensitivity to the shear layer thickness; however, we only considered two similar values, $d_s/H = 0.06$ and 0.08 , to ensure the current was negligible at the crest of the ridge. Strengthening or deepening the current increased the asymmetries.

R , the ratio of the magnitude of the downstream and upstream energy fluxes, varied over a wide range of values. For each current this ratio increased monotonically as the ridge width increased, being considerably less than 1 for the narrowest ridge ($a/H = 1.2$) and much larger than 1 for the widest ridge ($a/H = 6$), the transition point with $R = 1$ being approximately 3.5. For the widest ridge using a current of depth $z_s/H = -0.3$, the ratio increased from approximately 2 to 8 as the strength of the currents increased from $U_s/NH = 0.04$ to 0.16 while for the narrowest ridge R decreased from 0.8 to 0.25. Deepening the current increased these changes: for the strongest, deepest current (set 8) $R = 0.2$ for the narrowest ridge and 18 for the widest ridge.

In addition to introducing asymmetries in the energy fluxes in the upstream and downstream directions, the current has a significant effect on relative contributions of the various parts of the total flux. First, there is a leading-order contribution from the kinetic and available potential energy fluxes. The contribution from the available potential energy flux is small while that from the kinetic energy flux is significant. Furthermore, splitting the kinetic energy flux into contributions that are linear and quadratic in the velocity perturbation fields (from integrals of $\bar{U}^2 u'$ and $\bar{U} u'^2$ respectively where u' is the horizontal velocity perturbation) shows that the former is also significant for $U_s/NH = 0.1$. This can only come from mean second-order fields. Similarly, splitting the mean work term into contributions that are linear and quadratic in the perturbation fields, i.e. from integrals of $\bar{U} p_d$ and $u' p_d$ where p_d is the pressure perturbation, shows that the term linear in the pressure perturbation also makes a significant contribution. For the case with $a/H = 1.2$ and $U_s/NH = 0.1$ shown in

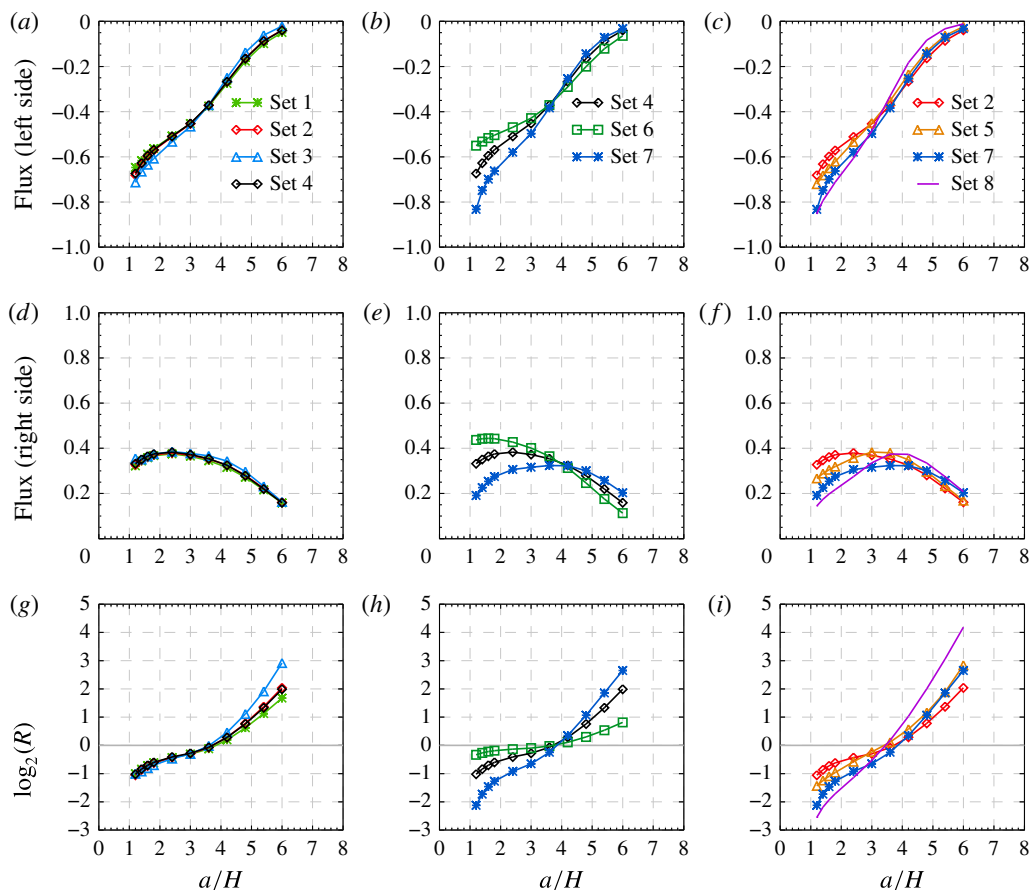


FIGURE 15. Tidally averaged energy fluxes at $t=20T$ scaled by F_0 as a function of the ridge width a/H for eight sets of simulations. (a,d,g) Sets 1–4 (legend in panel a). In sets 1–3 the ridge amplitude is varied. Sets 2 and 4 have different shear layer thicknesses. (b,e,h) Sets 4, 6, 7 (legend in panel b) vary primarily in the strength of the current ($U_s/(NH) = 0.1, 0.04$ and 0.16 respectively). (c,f,i) Sets 2, 5, 7 and 8 (legend in panel c). Sets 2, 5 differ only in depth of the current as do sets 7 and 8. These two pairs of sets differ in the strength of the current. The first two rows show fluxes to the left and right of the ridge (at $\xi = 3a/H$ and $-3a/H$). The third row shows the absolute value of the ratio of the rightward to leftward fluxes (log base 2).

figure 12 the contribution from the terms linear in the perturbations is approximately 25 % of the total leftward energy flux.

The tidally averaged mean fields show strong asymmetries about narrow ridges (figure 2) and are responsible for the tidally averaged work and kinetic energy flux terms undergoing significant quasi-periodic oscillations as a function of distance from the ridge. We leave further exploration of the mean currents to future work but note that the generation of mean currents when an internal wave beam impinges on a pycnocline has been observed in the laboratory (Mercier *et al.* 2012).

A linear theory, valid for small amplitude ridges and small tidal excursions, was developed to predict the leading-order perturbation fields. The fluxes based on the first-order solution (which does not include fluxes associated with second-order mean

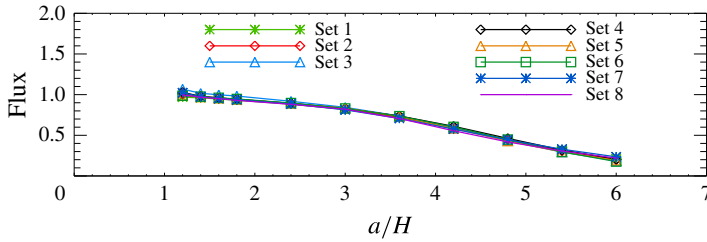


FIGURE 16. Total tidally averaged energy fluxes $\langle E_f \rangle|_{\xi_l}^{\xi_r}$ radiating away from the ridge as a function of the ridge width. Values are at $t=20T$ and are scaled by F_0 . $\xi_r = 3a/H$ and $\xi_l = -3a/H$.

fields which are formally just as important) predicts spatially varying work and kinetic energy flux terms that agree well with the second-order fluxes from the nonlinear simulations for narrow ridges (see figure 12). These spatial variations arise because in the presence of a background current the linear eigenmodes are not orthogonal. Both the amplitude of the spatial oscillations and their wavelength seen in the nonlinear simulations is well predicted by the first-order linear theory. For the case illustrated in figure 12 the wavelength of the oscillations is equal to the distance between locations where beams reflect off the surface and is close to the wavelength associated with interactions of the mode-one and mode-two waves. For a much wider ridge ($a/H = 6$, figure 13) the agreement between first-order linear theory and the second-order perturbation fluxes is not as good but improves as the ridge amplitude decreases. This may be because as the ridge width increases waves above the ridge propagate for longer distances in slightly shallower water, an effect not taken into account by linear theory.

We have considered an idealized situation in a nearly linear regime. For more realistic stratifications with a strong subsurface pycnocline wave propagation speeds would be larger. For the same currents used here this would result in smaller Froude numbers presumably reducing the strength of the asymmetry in the upstream and downstream fluxes. This will be explored in future work. Other complicating factors include three-dimensional time varying background currents, multiple tidal constituents leading to spring-neap tidal cycles, three-dimensional bathymetry and horizontally varying stratifications. In addition rotational effects should be considered. It would be extremely difficult to separate any observed mean fields into a background state and wave-induced mean fields unless the wave-induced mean fields were much stronger than in our simulations as may potentially occur in shallower water with stronger tidal currents than those used here. Nevertheless the findings reported here suggest that in the presence of strong currents internal wave energy fluxes emanating away from a quasi-two-dimensional ridge can be strongly asymmetric with the direction of the largest energy fluxes being sensitive to the ridge width. The results also suggest that estimating the traditional work term $\int u'p' dz$ may not be sufficient: it may undergo significant spatial variations and contributions from the kinetic energy fluxes need to be considered. It should be possible to measure the latter and see if it is a significant fraction of the work term $\langle W \rangle$ which is traditionally measured.

There are many possible avenues for future study. These include situations in which the bathymetry penetrates into the current as well as the effects of background currents on internal solitary wave generation when the stratification is non-constant, supercritical topography and three-dimensional effects. A full second-order theory

which includes fluxes associated with second-order mean fields is desirable as is one not restricted to small amplitude topography. A possible avenue for the latter, which is restricted to small excursion distances, is a Green's function approach similar to that developed for tide–topography interactions in the absence of a background current (Mathur, Carter & Peacock 2016). The theory could also be extended to include a free surface.

In conclusion, the presence of a uni-directional surface trapped background current introduces asymmetries in the internal wave field and in the energy fluxes radiating upstream and downstream while the total energy flux away from the ridge is not significantly affected. This would likely not be the case if the ridge penetrated into the current. For narrow ridges the upstream fluxes can be 4–5 times larger than the downstream fluxes while for sufficiently wide ridges the downstream fluxes may be 8 or more times larger than the upstream fluxes. In addition the kinetic energy flux can be significant as can contributions from second-order mean fields.

Acknowledgements

This work is supported by grants from the Natural Sciences and Engineering Research Council of Canada and from the Canadian Foundation for Innovation and the Ontario Research Fund. The facilities of the Shared Hierarchical Academic Research Computing Network (SHARCNET: <https://www.sharcnet.ca>) and Compute/Calcul Canada were used for this work.

REFERENCES

- AIKI, H., MATTHEWS, J. P. & LAMB, K. G. 2011 Modeling and energetics of tidally generated wave trains in the Lombok Strait: impact of the Indonesian throughflow. *J. Geophys. Res.* **116**, C03023.
- BELL, J. B., COLELLA, P. & GLAZ, H. M. 1989 A second-order projection method for the incompressible Navier–Stokes equations. *J. Comput. Phys.* **85**, 257–283.
- BELL, T. H. 1975*a* Lee waves in stratified flow with simple harmonic time dependence. *J. Fluid Mech.* **67**, 705–722.
- BELL, T. H. 1975*b* Topographically generated internal waves in the open ocean. *J. Geophys. Res.* **80**, 320–327.
- BROUTMAN, D., ROTTMAN, J. W. & ECKERMANN, S. D. 2004 Ray methods for internal waves in the atmosphere and ocean. *Annu. Rev. Fluid Mech.* **36**, 233–253.
- BUJISMAN, M. C., ANSONG, J. K., ARBIC, B. K., RICHMAN, J. G., SHRIVER, J. F., TIMKO, P. G., WALLCRAFT, A. J., WHALEN, C. B. & ZHAO, Z.-X. 2016 Impact of parameterized internal wave drag on the semidiurnal energy balance in a global ocean circulation model. *J. Phys. Oceanogr.* **46**, 1399–1419.
- EGBERT, G. D. & RAY, R. D. 2001 Estimates of M_2 tidal energy dissipation from TOPEX/Poseidon altimeter data. *J. Geophys. Oceanogr.* **106**, 24 821–24 852.
- FALAHAT, S., NYCANDER, J., ROQUET, F., THURNHERR, A. M. & HIBIYA, T. 2014 Comparison of calculated energy flux of internal tides with microstructure measurements. *Tellus A* **66**, 23240.
- FERRARI, R., MASHAYEK, A., MCDUGALL, T. J., NIKURASHIN, M. & CAMPIN, J.-M. 2016 Turning ocean mixing upside down. *J. Phys. Oceanogr.* **46**, 2239–2261.
- GARRETT, C. & KUNZE, E. 2007 Internal tide generation in the deep ocean. *Annu. Rev. Fluid Mech.* **39**, 57–87.
- HOLLIDAY, D. & MCINTYRE, M. E. 1981 On potential energy density in an incompressible, stratified fluid. *J. Fluid Mech.* **107**, 221–225.
- KELLY, S. M., NASH, J. D. & KUNZE, E. 2010 Internal-tide energy over topography. *J. Geophys. Res.* **115**, C06014.

- KHATIWALA, S. 2003 Generation of internal tides in an ocean of finite depth: analytical and numerical calculations. *Deep-Sea Res.* I **50**, 3–21.
- LAMB, K. G. 1994 Numerical experiments of internal wave generation by strong tidal flow across a finite amplitude bank edge. *J. Geophys. Res.* **99**, 843–864.
- LAMB, K. G. 2007 Energy and pseudoenergy flux in the internal wave field generated by tidal flow over topography. *Cont. Shelf Res.* **27**, 1208–1232.
- LAMB, K. G. 2008 On the calculation of the available potential energy of an isolated perturbation in a density-stratified fluid. *J. Fluid Mech.* **597**, 415–427.
- LAMB, K. G. 2010 Energetics of internal solitary waves in a background sheared current. *Nonlinear Process. Geophys.* **17**, 553–568.
- LAMB, K. G. & NGUYEN, V. T. 2009 Calculating energy flux in internal solitary waves with an application to reflectance. *J. Phys. Oceanogr.* **39**, 559–580.
- LEFAUVE, A., MULLER, C. & MELET, A. 2015 A three-dimensional map of tidal dissipation over abyssal hills. *J. Geophys. Res. Oceans* **120**, 4760–4777.
- LLEWELLYN SMITH, S. G. & YOUNG, W. R. 2002 Conversion of the barotropic tide. *J. Phys. Oceanogr.* **32**, 1544–1566.
- LYARD, F., LEFEVRE, F., LETELLIER, T. & FRANCIS, O. 2006 Modelling the global ocean tides: modern insights from FES2004/2004. *Ocean Dyn.* **56**, 394–415.
- MATHUR, M., CARTER, G. S. & PEACOCK, T. 2016 Internal tide generation using Green function analysis: to WKB or not to WKB. *J. Phys. Oceanogr.* **46**, 2157–2168.
- MATTHEWS, J. P., AIKI, H., MASUDA, S., AWAJI, T. & ISHIKAWA, Y. 2011 Monsoon regulation of Lombok Strait internal waves. *J. Geophys. Res.* **116**, C05007.
- MERCIER, M. J., MATHUR, M., GOSTIAUX, L., GERKEMA, T., MAGALHÃES, J. M., DA SILVA, J. C. B. & DAUXOIS, T. 2012 Soliton generation by internal tidal beams impinging on a pycnocline: laboratory experiments. *J. Fluid Mech.* **704**, 37–60.
- MUNK, W. & WUNSCH, C. 1998 Abyssal recipes. Part II. Energetics of tidal and wind mixing. *Deep-Sea Res.* **45**, 1977–2010.
- NASH, J. D., ALFORD, M. H. & KUNZE, E. 2005 Estimating internal wave energy fluxes in the ocean. *J. Atmos. Ocean. Technol.* **22**, 1551–1570.
- NIKURASHIN, M. & FERRARI, R. 2013 Overturning circulation driven by breaking internal waves in the deep ocean. *Geophys. Res. Lett.* **40**, 3133–3137.
- PICKERING, A., ALFORD, M., NASH, J., RAINVILLE, L., BUIJSMAN, M., KO, D. S. & LIM, B. 2015 Structure and variability of internal tides in Luzon Strait. *J. Phys. Oceanogr.* **45**, 1574–1594.
- POLZIN, K. L., TOOLE, J. M., LEDWELL, J. R. & SCHMITT, R. W. 1997 Spatial variability of turbulent mixing in the abyssal ocean. *Science* **276**, 93–96.
- SCOTT, R. B., GOFF, J. A., NAVEIRA GARABATO, A. C. & NURSER, A. J. G. 2011 Global rate and spectral characteristics of internal gravity wave generation by geostrophic flow over topography. *J. Geophys. Res.* **116**, C09029.
- SCOTTI, A., BEARDSLEY, R. & BUTMAN, B. 2006 On the interpretation of energy and energy fluxes of nonlinear internal waves: an example from Massachusetts Bay. *J. Fluid Mech.* **561**, 103–112.
- SHEPHERD, T. G. 1993 A unified theory of available potential energy. *Atmos.-Ocean* **31**, 1–26.
- DA SILVA, J. C. B., NEW, A. I. & MAGALHAES, J. M. 2011 On the structure and propagation of internal solitary waves generated at the Mascarene plateau in the Indian ocean. *Deep-Sea Res.* I **58**, 229–240.
- WATERHOUSE, A. F., MACKINNON, J. A., NASH, J. D., ALFORD, M. H., KUNZE, E., SIMMONS, H. L., POLZIN, K. L., ST. LAURENT, L. C., SUN, O. M., PINKEL, R. *et al.* 2014 Global patterns of diapycnal mixing from measurements of the turbulent dissipation rate. *J. Phys. Oceanogr.* **44**, 1854–1872.
- WUNSCH, C. & FERRARI, R. 2004 Vertical mixing, energy, and the general circulation of the oceans. *Annu. Rev. Fluid Mech.* **36**, 281–314.